The figure indicates how similar different GRBs are to each other. Points which are closer together are more similar, and points which are further away are more different. What we find is that there are two distinct groups, one orange and the other blue. The orange dots appear to correspond to "short" GRB, which have been hypothesized to be produced by mergers of neutron stars, and the blue dots appear to correspond to "long" GRB, which might instead be produced by the collapse of dying, massive stars. Credit: Niels Bohr Institute

By applying a machine-learning algorithm, scientists at the Niels Bohr Institute, University of Copenhagen, have developed a method to classify all gamma-ray bursts (GRBs), rapid highly energetic explosions in distant galaxies, without needing to find an afterglow—by which GRBs are presently categorized. This breakthrough, initiated by first-year B.Sc. students, may prove key in finally discovering the origins of these mysterious bursts. The result is now published in Astrophysical Journal Letters.

Breakthrough achieved using machine-learning algorithm

Now, scientists at the Niels Bohr Institute have developed a method to classify all GRBs without needing to find an afterglow. The group, led by first-year B.Sc. Physics students Johann Bock Severin, Christian Kragh Jespersen and Jonas Vinther, applied a machine-learning algorithm to classify GRBs. They identified a clean separation between long and short GRB's. Their work, carried out under the supervision of Charles Steinhardt, will bring astronomers a step closer to understanding GRB's.

Ever since gamma-ray bursts (GRBs) were accidentally picked up by Cold War satellites in the 70s, the origin of these rapid bursts have been a significant puzzle. Although many astronomers agree that GRBs can be divided into shorter (typically less than 1 second) and longer (up to a few minutes) bursts, the two groups are overlapping. It has been thought that longer bursts might be associated with the collapse of massive stars, while shorter bursts might instead be caused by the merger of neutron stars. However, without the ability to separate the two groups and pinpoint their properties, it has been impossible to test these ideas.

So far, it has only been possible to determine the type of a GRB about 1% of the time, when a telescope was able to point at the burst location quickly enough to pick up residual light, called an afterglow. This has been such a crucial step that astronomers have developed worldwide networks capable of interrupting other work and repointing large telescopes within minutes of the discovery of a new burst. One GRB was even detected by the LIGO Observatory using gravitational waves, for which the team was awarded the 2017 Nobel Prize.
Shining light on the data

The preparation of the feature space—the input you give the algorithm—was the most challenging part of the project, says Johann Bock Severin. Essentially, the students had to prepare the dataset in such a way that its most important features would stand out. "I like to compare it to hanging your data points from the ceiling in a dark room," explains Christian Kragh Jespersen. "Our main problem was to figure out from what direction we should shine light on the data to make the separations visible."

'Step 0 in understanding GRB's'

The students explored the t-SNE machine-learning algorithm as part of their 1st Year project, a 1st year course in the Bachelor of Physics. "By the time we got to the end of the course, it was clear we had quite a significant result", their supervisor Charles Steinhardt says. The students' mapping of the t-SNE cleanly divides all GRB's from the Swift observatory into two groups. Importantly, it classifies GRB's that previously were difficult to classify. "This essentially is step 0 in understanding GRB's," explains Steinhardt. "For the first time, we can confirm that shorter and longer GRB's are indeed completely separate things."

Without any prior theoretical background in astronomy, the students have discovered a key piece of the puzzle surrounding GRB's. From here, astronomers can start to develop models to identify the characteristics of these two separate classes.


Provided by Niels Bohr Institute