

Tarantula Nebula spins web of mystery in Spitzer image

27 January 2020, by Calla Cofield



This image from NASA's Spitzer Space Telescope shows the Tarantula Nebula in two wavelengths of infrared light. The red regions indicate the presence of particularly hot gas, while the blue regions are interstellar dust that is similar in composition to ash from coal or wood-burning fires on Earth. Credit: NASA/JPL-Caltech

The Tarantula Nebula, seen in this image by the Spitzer Space Telescope, was one of the first targets studied by the infrared observatory after its launch in 2003, and the telescope has revisited it many times since. Now that Spitzer is set to be retired on Jan. 30, 2020, scientists have generated a new view of the nebula from Spitzer data.

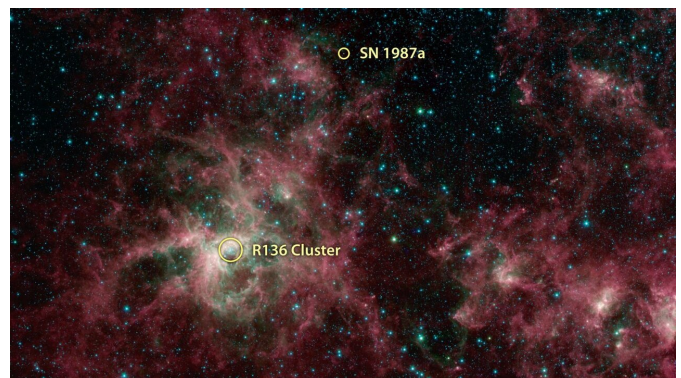
This high-resolution image combines data from multiple Spitzer observations, most recently in February and September 2019.

"I think we chose the Tarantula Nebula as one of our first targets because we knew it would demonstrate the breadth of Spitzer's capabilities," said Michael Werner, who has been Spitzer's project scientist since the mission's inception and is based at NASA's Jet Propulsion Laboratory in Pasadena, California. "That region has a lot of interesting dust structures and a lot of star formation happening, and those are both areas

where infrared observatories can see a lot of things that you can't see in other wavelengths."

Infrared light is invisible to the human eye, but some wavelengths of infrared can pass through clouds of gas and dust where [visible light](#) cannot. So scientists use infrared observations to view [newborn stars](#) and still-forming "protostars," swaddled in the clouds of gas and dust from which they formed.

Located in the Large Magellanic Cloud—a dwarf galaxy gravitationally bound to our Milky Way galaxy—the Tarantula Nebula is a hotbed of star formation. In the case of the Large Magellanic Cloud, such studies have helped scientists learn about rates of star formation in galaxies other than the Milky Way.



This annotated image from NASA's Spitzer Space Telescope shows the Tarantula Nebula in infrared light. The supernova 1987A and the starburst region R136 are noted. The magenta-colored regions are primarily interstellar dust that is similar in composition to ash from coal or wood fires on Earth. Credit: NASA/JPL-Caltech

The [nebula](#) also hosts R136, a "starburst" region, where massive stars form in extremely close proximity and at a rate far higher than in the rest of

the galaxy. Within R136, in an area less than 1 light-year across (about 6 trillion miles, or 9 trillion kilometers), there are more than 40 massive stars, each containing at least 50 times the mass of our Sun. By contrast, there are no stars at all within 1 light-year of our Sun. Similar starburst regions have been found in other galaxies, containing dozens of massive stars—a higher number of [massive stars](#) than what is typically found in the rest of their host galaxies. How these starburst regions arise remains a mystery.

On the outskirts of the Tarantula Nebula, you can also find one of astronomy's most-studied [stars](#) that has exploded in a supernova. Dubbed 1987A because it was the first supernova spotted in 1987, the exploded star burned with the power of 100 million Suns for months. The shockwave from that event continues to move outward into space, encountering material ejected from the star during its dramatic death.

When the shockwave collides with dust, the dust heats up and begins to radiate in infrared light. In 2006, Spitzer observations saw that light and determined that the dust is largely composed of silicates, a key ingredient in the formation of rocky planets in our solar system. In 2019, scientists used Spitzer to study 1987A to monitor the changing brightness of the expanding shockwave and debris to learn more about how these explosions change their surrounding environment.

Provided by Jet Propulsion Laboratory

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