

Quest for source of black hole dark matter

14 August 2018, by Anne M Stark



LLNL scientists Michael Schneider, Will Dawson, Nathan Golovich and George Chapline look for black holes in the Lab's telescope remote observing room. Credit: Julie Russell/LLNL.

Like a game of "hide and seek," Lawrence Livermore astrophysicists know that there are black holes hiding in the Milky Way, just not where.

If they find them toward the galactic bulge (a tightly packed group of [stars](#)) and the Magellanic Clouds, then [black holes](#) as massive as 10,000 times the mass of the sun might make up [dark matter](#). If they are only toward the [galactic bulge](#) then they are probably just from a few dead stars.

Typically, to observe the Magellanic Clouds, scientists must travel to observatories in the Southern Hemisphere.

But recently, the LLNL team got a new tool that's a little closer to home to help them in the search. As part of the Space Science and Security Program and an LDRD project, LLNL has a new telescope remote observing room.

The team is using the observing room to conduct a gravitational [microlensing](#) survey of the Milky Way and Magellanic Clouds in search of intermediate mass black holes (approximately 10 to 10,000

times the mass of the sun) that may make up the majority of dark [matter](#).

"The remote observing room enables us to control the National Optical Astronomers Observatory Blanco 4-meter telescope located in Chile at the Cerro Tololo Inter-American Observatory," said LLNL principal investigator Will Dawson. The team already has conducted its first observing run with the remote observing room.

The visible universe is composed of approximately 70 percent dark energy, 25 percent dark matter and 5 percent normal matter. However, dark matter has remained a mystery since it was first postulated in 1933. The MACHO Survey, led by Lawrence Livermore in the 1990s, sought to test whether dark matter was composed of baryonic massive compact halo objects (MACHOs). The survey concluded that baryonic MACHOs smaller than 10 solar masses could not account for more than 40 percent of the total dark matter mass.

Recently, the discovery of two merging black holes has renewed interest in MACHO dark matter composed of primordial black holes (formed in the early universe, before the first stars) with approximately 10 to 10,000 solar masses. This is an idea first proposed in 1975 by LLNL physicist and project co-investigator George Chapline. The most direct means of exploring this mass range is by searching for the gravitational microlensing signal in existing archival astronomical imaging and carrying out a next-generation microlensing survey with state-of-the-art wide-field optical imagers on telescopes 10 to 25 times more powerful than those used in the original MACHO surveys.

Microlensing is an astronomical effect predicted by Einstein's general theory of relativity. According to Einstein, when the light emanating from a star passes very close to another massive object (e.g., black hole) on its way to an observer on Earth, the gravity of the intermediary massive object will slightly bend and focus the light rays from the source star, causing the lensed background star to

appear brighter than it normally would.

"We are developing a novel means of microlensing detection that will enable us to detect the parallactic microlensing signature associated with black holes in this mass range," Dawson said. "We will detect and constrain the fraction of dark matter composed of intermediate mass black holes and measure their [mass](#) spectrum in the Milky Way."

Provided by Lawrence Livermore National Laboratory

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