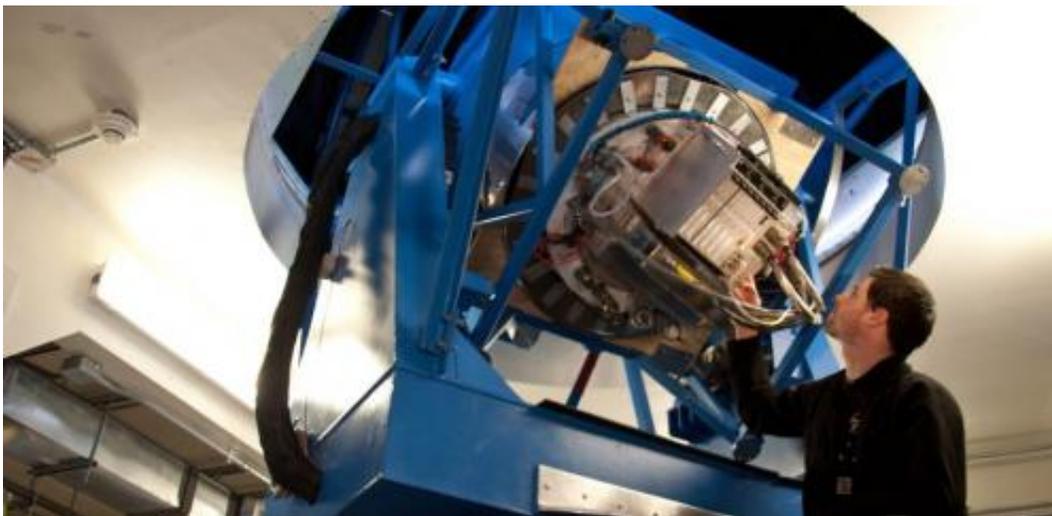


First hints of gravitational waves in the Big Bang's afterglow

March 18 2014, by Krzysztof Bolejko



Graduate student Justus Brevik testing the BICEP2 used to find evidence of cosmic inflation nearly 14 billion years ago. Credit: EPA/Steffen Richter/Harvard University

Scientists at the Harvard-Smithsonian Centre for Astrophysics in the US have [announced](#) overnight what they believe is the indirect detection of gravitational waves in the afterglow of the Big Bang.

The [discovery](#) by the Background Imaging of Cosmic Extragalactic Polarisation ([BICEP](#)) collaboration, indeed even [the rumours of such a discovery](#), sparked a huge discussion among the scientific community. Why?

As the last untested prediction of Einstein's Theory of General Relativity, finding [gravitational waves](#) is a big deal.

The BICEP discovery provides further indirect evidence for the existence of gravitational waves (the 1993 Nobel Prize in Physics was awarded to [Russell Hulse and Joseph Taylor](#) for finding a double pulsar that strongly supported these "ripples" in spacetime).

Secondly, and most importantly, it advances our knowledge of the [universe](#) enormously.

Before this announcement, thanks to [Big Bang nucleosynthesis](#) (when light elements such as hydrogen and helium were created), we could measure the universe back to about a minute after the Big Bang.

The finding today has allowed us to study the universe when it was a trillionth of a trillionth of a trillionth of a second old, when so-called "[inflation](#)" took place.

Inflation was a period of accelerated expansion of the early universe, but before we explain what that is and why it's so important, first a few words about what was actually detected by the [BICEP](#) telescope.

Cosmic microwave background

Everywhere astronomers point their (microwave) telescopes there is a faint glow of light. The picture formed from this (microwave) light is almost perfectly the same in all directions and corresponds to an object shining with a temperature of just 2.72548 Kelvin above absolute zero (or a chilly -270.42452 Celsius).

This afterglow of the Big Bang, called the [cosmic microwave background](#) ([CMB](#)), is a fossil radiation emitted when the universe was

still very young, just 380,000 years old.

Why the CMB looks so similar in all directions was one of the greatest riddles of 20th century cosmology.

Another conundrum that puzzled astronomers was the source of tiny changes in temperature that were discovered by the [COBE](#) satellite (winning [the 2006 Nobel Prize](#)), and appearing as blue and red spots in the picture above from the [WMAP satellite](#). Inflation provides a neat solution to these problems.

Inflation

So what does inflation mean for the early universe? Generically, inflation theories suggest the universe expanded from a tiny 10^{-33} cm (the [Planck scale](#)) to many times larger than the visible universe today ($\gg 10^{28}$ cm).

This means that a small region of the young universe with the same temperature is expanded to larger than the entire observable universe, explaining why the observed CMB has such a similar temperature everywhere.

Furthermore, the world of the very small (governed by quantum mechanics) has been expanded into the realm of the very large. One thing that Quantum Mechanics tells us is that on the smallest scales everything fluctuates.



The BICEP2 telescope at the Dark Sector Lab at Amundsen-Scott South Pole Station. Credit: Harvard University

After inflation these tiny fluctuations are expanded to enormous scales, becoming the imperfections we see in the CMB. These are the gravitational seeds around which galaxies will later form.

The problem of why the CMB is so similar everywhere but not *perfectly* so is naturally answered by inflation.

There is another signature that inflation predicts. Fluctuations in not just what would become the normal matter we see in the CMB, but in a background gravitational field which delicately imprints itself on the light from the CMB.

This can't be seen in the normal maps of the CMB (such as the above

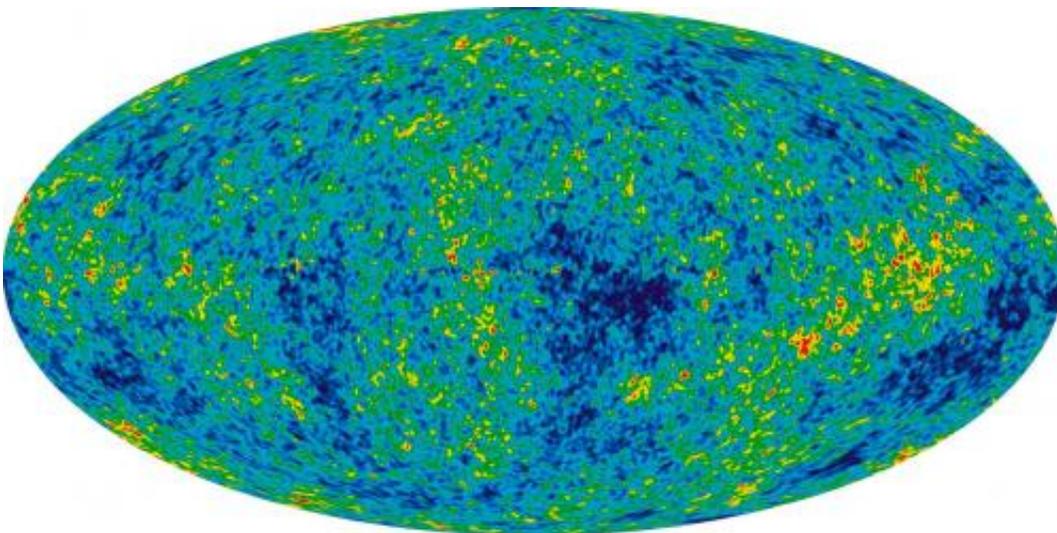
picture), but instead in the polarisation (the direction in which light oscillates as it travels) of the CMB.

Polarisation of the CMB

The BICEP telescope split the polarised CMB light into two types of shapes or "modes": B-modes and E-modes.

It is relatively easy to create the E-modes and this pattern of polarisation has been known since 2002 with the [DASI telescope](#).

Much harder to create (and a far weaker signal) is the B-mode. These can only come from gravitational lensing of the CMB light by intervening galaxies or gravitational waves from the [early universe](#) expanded to enormous scales by inflation. The B-mode polarisation induced by [gravitational lensing](#) was detected [last year](#). Today it's the turn of the gravitational waves.



Cosmic Microwave Background (CMB) radiation. The blue spots correspond to colder (approximately 2.7253 K or -270.4247 °C), while red spots to warmer regions (approximately 2.7257 K or -270.4243 °C). Credit: NASA / WMAP

Science Team

Gravitational waves from the early universe

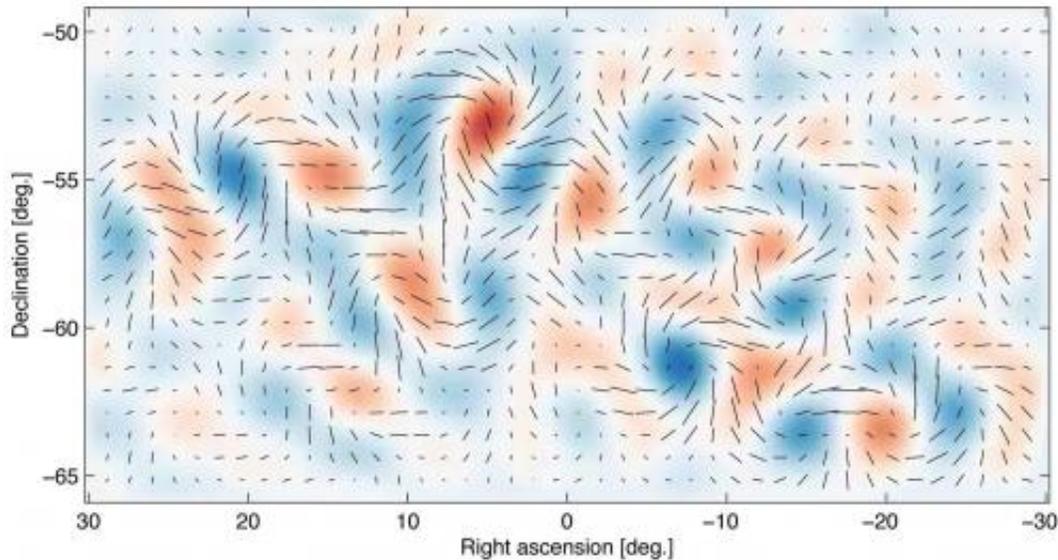
Alternative theories to inflation do not produce gravitational waves so the existence of B-modes detected by BICEP is strong evidence not only of the gravitational wave background but also inflation itself.

The B-mode signal of the gravitational waves is a bit stronger than previous efforts such as the Planck satellite [had suggested](#).

This result implies the energy scales at which inflation kicks in is close to that of Grand Unified Theory, meaning that inflation could occur even sooner after the Big Bang.

It is also the first indirect detection of the gravitational wave *background*. So far there has still been no direct detection of the gravitational radiation.

The first direct detection should follow in a few months, when the Advanced Laser Interferometer Gravitational-Wave Observatory (or Advanced [LIGO](#) for short) will start to operate. It is envisaged that the experiment will directly detect [gravitational radiation](#) coming from astrophysical sources from nearby galaxies.



Gravitational waves from inflation generate a faint but distinctive twisting pattern in the polarisation of the cosmic microwave background, known as a “curl” or B-mode pattern. The red and blue shading shows the degree of twisting of this B-mode pattern. Credit: EPA/Harvard University

Gravitational insight into the origin of the universe

If today's announcement by BICEP is confirmed by other experiments it will be a huge boost to the theory of inflation and the existence of gravitational waves.

Everything we know about the world around us is based on seeing light (electromagnetic waves). Yet this discovery opens up the possibility of an entirely new sense with which to view our universe – gravitational waves.

Who knows what they will allow us to see? At the very least we will be able to "see" into the hearts of exploding stars, or titanic collisions between galaxies and of course right back to the start of our entire

universe as shown by this discovery.

It will allow us access to the unimaginable world of our universe just after the Big Bang and has already given hints of a new realm of physics, perhaps a more significant discovery for particle physics than even the confirmation of the Higgs Boson.

More accurate measurements of the gravitational waves across different scales on the sky will allow us to test models of inflation. All of which has taken us one step closer to answering the ultimate question of the nature of the Big Bang itself.

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