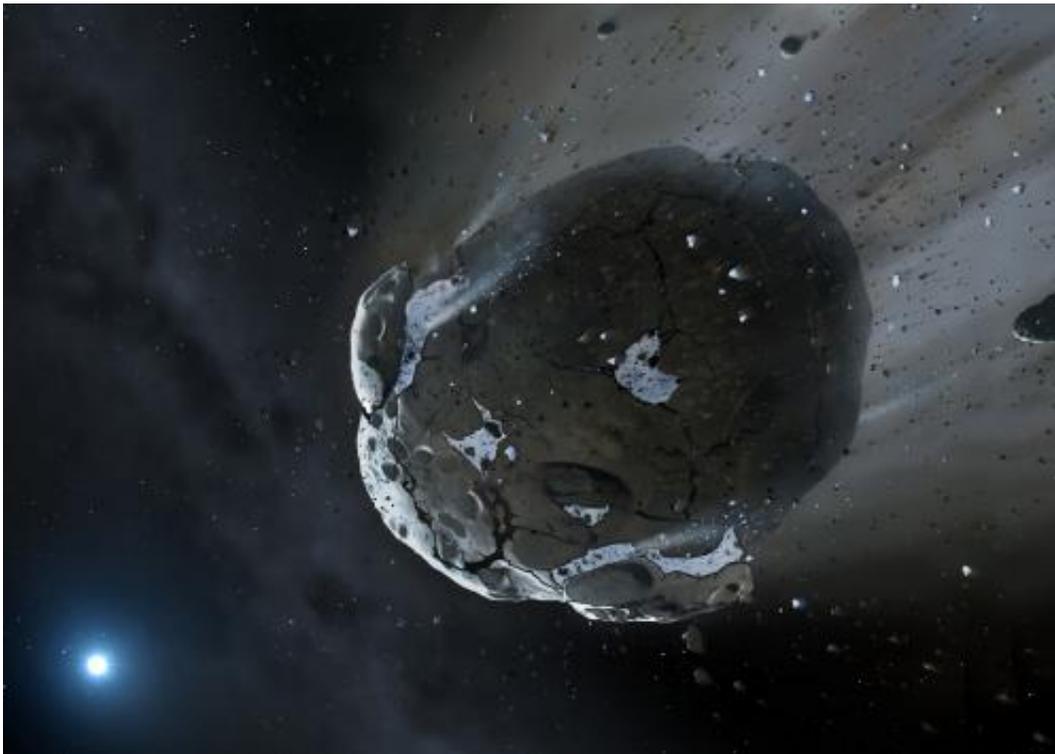


Watery asteroid discovered in dying star points to habitable exoplanets

October 10 2013



Artist's impression of a rocky and water-rich asteroid being torn apart by the strong gravity of the white dwarf star GD 61. Similar objects in the solar system likely delivered the bulk of water on Earth and represent the building blocks of the terrestrial planets. Credit: © Mark A. Garlick, space-art.co.uk, University of Warwick and University of Cambridge

Astronomers have found the shattered remains of an asteroid that contained huge amounts of water orbiting an exhausted star, or white

dwarf. This suggests that the star GD 61 and its planetary system – located about 150 light years away and at the end of its life – had the potential to contain Earth-like exoplanets, they say.

This is the first time that both water and a [rocky surface](#) - two "key ingredients" for [habitable planets](#) - have been found together beyond our [solar system](#).

Earth is essentially a 'dry' planet, with only 0.02% of its mass as surface water, so oceans came long after it had formed; most likely when water-rich asteroids in the solar system crashed into our planet.

The new discovery shows that the same water 'delivery system' could have occurred in this distant, dying star's solar system – as latest evidence points to it containing a similar type of water-rich asteroid that would have first brought water to Earth.

The asteroid analysed is composed of 26% water mass, very similar to Ceres, the largest asteroid in the main belt of our solar system. Both are vastly more water-rich compared with Earth.

Astronomers at the Universities of Cambridge and Warwick say this is the first "reliable evidence" for water-rich, rocky planetary material in any extrasolar planetary system.

They describe it as a "look into our future" as, six billion years from now, alien [astronomers](#) studying the rocky [remains](#) around our burned out sun might reach the same conclusion - that terrestrial [planets](#) once circled our parent star.

The new research findings used NASA's Hubble Space Telescope and are reported today in the journal *Science*.

All rocky planets form from the accumulation of asteroids, growing until full size, so asteroids are essentially the 'building blocks' of planets.

"The finding of water in a large asteroid means the building blocks of habitable planets existed – and maybe still exist – in the GD 61 system, and likely also around substantial number of similar parent stars," said lead author Jay Farihi, from Cambridge's Institute of Astronomy.

"These water-rich building blocks, and the terrestrial planets they build, may in fact be common – a system cannot create things as big as asteroids and avoid building planets, and GD 61 had the ingredients to deliver lots of water to their surfaces," Farihi said.

"Our results demonstrate that there was definitely potential for habitable planets in this exoplanetary system."

The researchers say that the water detected most likely came from a minor planet, at least 90 km in diameter but probably much larger, that once orbited the GD 61 star before it became a white dwarf around 200 million years ago.

Previous and current astronomical observations have measured the size and density of exoplanets, but not their composition.

This is because conventional work was done on planets orbiting living stars. But the only way to see what a distant planet is made of is to take it apart, say the researchers, and nature does this for us in a dying white dwarf system through its extreme gravitational pull – sucking in and shredding the surrounding material.

This debris, which "pollutes" the atmosphere of the white dwarf, can then be chemically analysed using powerful spectrograph techniques that "distill the entire asteroid, core and all", they say.

The team detected a range of "elemental abundance" in the white dwarf's contaminated atmosphere – such as magnesium, silicon and iron, which, together with oxygen are the main components of rocks.

By calculating the number of these elements relative to oxygen, the researchers were able to predict how much oxygen should be in the atmosphere of the white dwarf – but they found "significantly" more oxygen than if there were only rocks.

"This oxygen excess can be carried by either water or carbon, and in this star there is virtually no carbon – indicating there must have been substantial water," said co-author Boris Gänsicke, from the University of Warwick.

"This also rules out comets, which are rich in both water and carbon compounds, so we knew we were looking at a rocky asteroid with substantial [water](#) content – perhaps in the form of subsurface ice – like the asteroids we know in our solar system such as Ceres," Gänsicke said.

Ultraviolet observations are the only way to obtain such precise measurement of oxygen levels in the white dwarf's debris – and that can only be carried out above the Earth's atmosphere.

The team used the Cosmic Origins Spectrograph onboard Hubble to get the data required, with chemical analysis computed by team member Detlev Koester from the University of Kiel.

The "planetary bodies" such as these asteroids that fall into and pollute this dying star – which, in its heyday, was three times 'heavier' than our sun – also reveal that giant exoplanets probably still exist in this remote and withering system.

"In order for the asteroids to pass sufficiently close to the white dwarf to

be shredded, then eaten, they must be perturbed from the asteroid belt – essentially pushed – by a massive object like a giant planet," added Farihi.

"These asteroids tell us that the GD 61 system had – or still has – rocky, terrestrial planets, and the way they pollute the white dwarf tells us that giant planets probably still exist there.

"This supports the idea that the star originally had a full complement of terrestrial planets, and probably gas giant planets, orbiting it – a complex system similar to our own."

More information: "Evidence for Water in the Rocky Debris of a Disrupted Extrasolar Minor Planet," by J. Farihi et al. *Science*, 2013.

Provided by University of Cambridge

Citation: Watery asteroid discovered in dying star points to habitable exoplanets (2013, October 10) retrieved 25 April 2024 from <https://phys.org/news/2013-10-watery-asteroid-dying-star-habitable.html>

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