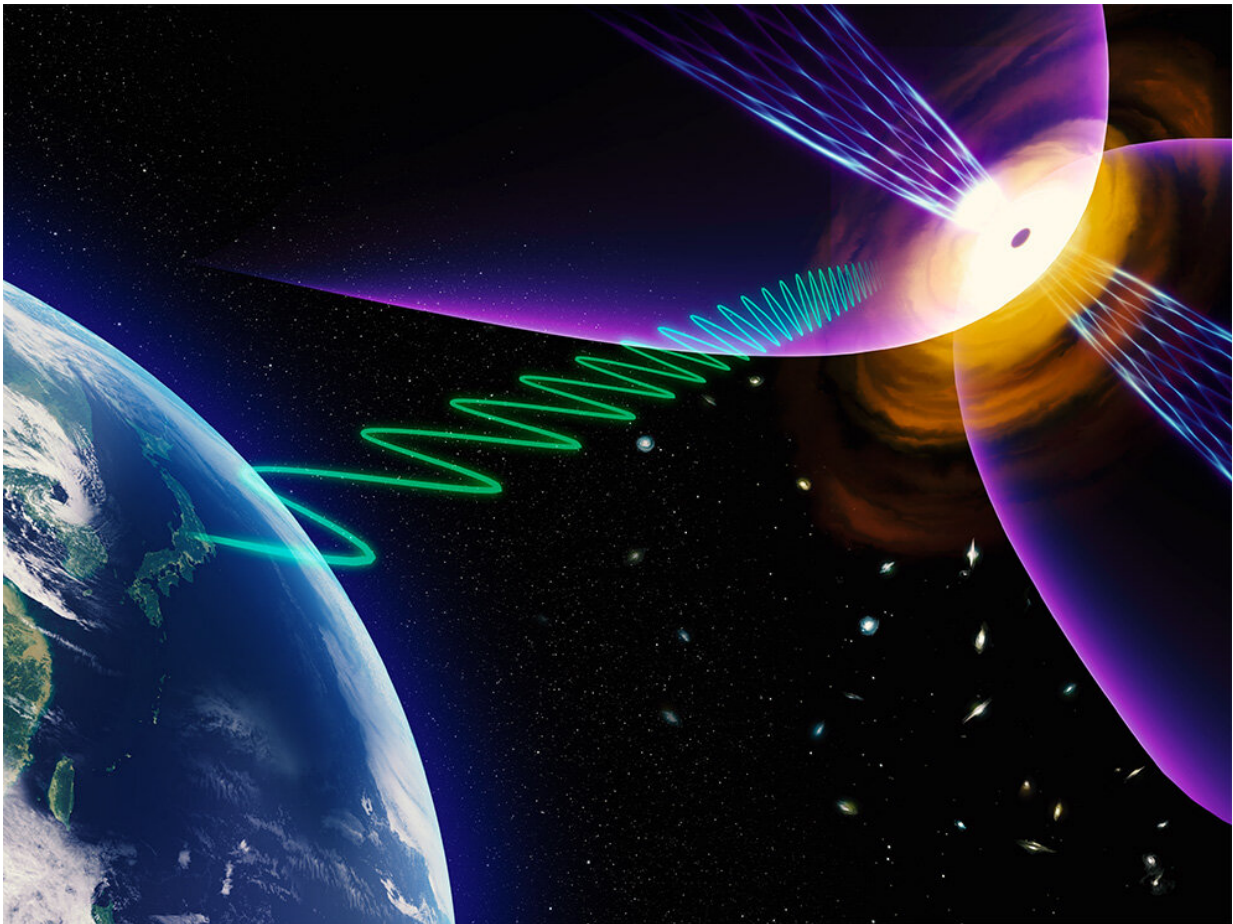


# VERA unveils surroundings of rapidly growing black holes

July 19 2023

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Jets ejected from a rapidly growing supermassive black hole with surrounding outflows. The polarization plane of a radio wave emitted from the vicinity of a black hole rotates as it passes through the surrounding magnetized gas. Credit: NAOJ

An international team of astronomers used the state-of-the-art capability of VERA, a Japanese network of radio telescopes operated by NAOJ, to uncover valuable clues about how rapidly growing "young" supermassive black holes form, grow, and possibly evolve into more powerful quasars.

It is now widely accepted that nearly every active galaxy harbors a [supermassive black hole](#) at its core, with masses ranging from millions to billions of times that of the sun. The growth history by which these [black holes](#) have gained such huge masses, however, remains an open question.

Led by Mieko Takamura, a graduate student at the University of Tokyo, an international team focused on a distinct category of active galaxies known as Narrow-line Seyfert 1 (NLS1) galaxies. These galaxies are suspected to contain relatively small yet rapidly growing [massive black holes](#), thus offering a potential opportunity to study an early evolutionary stage of these cosmic monsters.

To gain a deeper understanding of the immediate surroundings of these peculiar black holes, the team observed the cores of six nearby active NLS1 galaxies using VERA—a radio telescope network with vision more than 100,000 times more powerful than the human eye. In particular, the team leveraged the newly enhanced ultra-wideband recording capability of VERA, enabling them to detect faint "polarized" [radio waves](#) emanating from the core of these galaxies with unprecedented precision.

These results are published in an article titled, "Probing the heart of active narrow-line Seyfert 1 galaxies with VERA wideband polarimetry," in *The Astrophysical Journal*.

A portion of radio waves emitted near [supermassive black holes](#) is known to exhibit polarization. As this polarized emission propagates through the magnetized gas surrounding the black hole, the plane of

polarization gradually rotates, causing an effect known as Faraday rotation. The extent of this rotation (at a given wavelength) is proportional to the gas density and the strength of the magnetic field within the propagating medium. Therefore, polarization and Faraday rotation provide valuable insights into the immediate environment surrounding a central black hole.

Together with the sharpest-ever view towards the cores of these galaxies, the new data have unveiled significantly greater Faraday rotation compared to measurements obtained towards older, more-massive, well-developed black holes. This indicates the presence of abundant gas in the nuclear regions of these [galaxies](#), facilitating the rapid growth of the central black holes.

**More information:** Mieko Takamura et al, Probing the Heart of Active Narrow-line Seyfert 1 Galaxies with VERA Wideband Polarimetry, *The Astrophysical Journal* (2023). [DOI: 10.3847/1538-4357/acd9a8](#)

Provided by National Astronomical Observatory of Japan

Citation: VERA unveils surroundings of rapidly growing black holes (2023, July 19) retrieved 28 April 2024 from <https://phys.org/news/2023-07-vera-unveils-rapidly-black-holes.html>

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