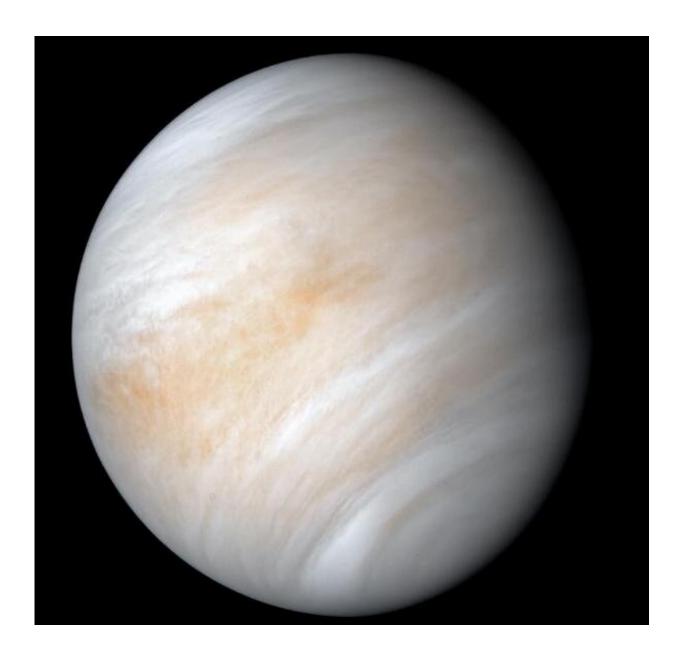


New analysis shows how sulfur clouds can form in Venus' atmosphere

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A reprocessed image of archived Mariner 10 data collected in 1974. This is a



false color image created using orange and ultraviolet filters for the red and blue channels, respectively. The clouds are at about 60 kilometers altitude, and the image illustrates the presence of an unknown ultraviolet absorber in the atmosphere, a long unsolved mystery of Venus. Credit: NASA/JPL-Caltech

Scientists using sophisticated computational chemistry techniques have identified a new pathway for how sulfur particles can form in the atmosphere of Venus. These results may help to understand the long sought-after identity of the mysterious ultraviolet absorber on Venus.

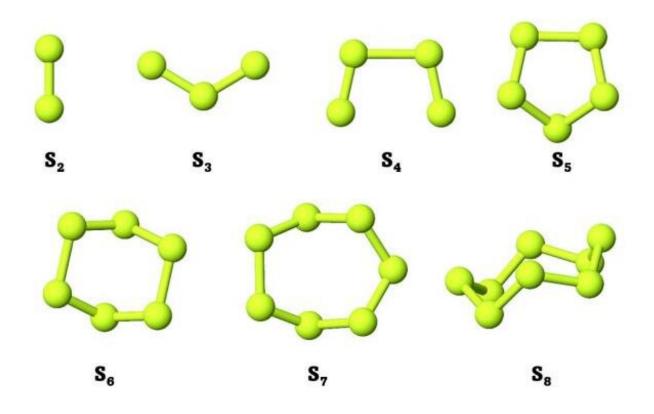
"We know that the <u>atmosphere</u> of Venus has abundant SO_2 and sulfuric acid particles. We expect that ultraviolet destruction of SO_2 produces <u>sulfur</u> particles. They are built up from atomic S (sulfur) to S_2 , then S_4 and finally S_8 . But how is this process initiated, that is, how does S_2 form?" said Planetary Science Institute Senior Scientist James Lyons, an author on the *Nature Communications* paper "Photochemical and thermochemical pathways to S_2 and polysulfur formation in the atmosphere of Venus."

One possibility is to form S_2 from two <u>sulfur atoms</u>, that is, reaction of S and S. Molecules of S_2 and S_2 can then combine to form S_4 , and so on. Sulfur particles can form either by condensation of S_8 or by condensation of S_2 , S_4 and other allotropes—different physical forms in which an element can exist—which then rearrange to form condensed S_8 .

"Sulfur particles, and the yellow sulfur we more commonly encounter, are made up of mostly S_8 , which has a ring structure. The <u>ring structure</u> makes S_8 more stable against destruction by UV light than the other allotropes. To form S_8 , we can either start with two S atoms and make S_2 , or we can produce S_2 by another pathway, which is what we've done



in the paper," said Lyons.



Sulfur molecules come in many forms called allotropes, from S_2 up to S_8 . The subscript indicates the number of S atoms in the allotrope. We are proposing here a new pathway to S_2 formation. With S_2 available in the atmosphere, S_4 and S_8 are produced. S_8 is the common form of yellow sulfur that can be seen near volcanic vents or that comes in a bottle. The sulfur allotropes S_3 and S_4 have been proposed to be the mysterious UV absorber in the Venus atmosphere. Although there is no consensus yet on the identity of the absorber, it's very likely that sulfur chemistry is involved. Credit: Figure adapted from Jackson et al., Chem. Sci., 2016, published by the Royal Society of Chemistry.

"We found a new pathway for S_2 formation, the reaction of sulfur monoxide (SO) and disulfur monoxide (S₂O), which is much faster than



combining two S atoms to make S₂," Lyons said.

"For the first time, we are using computational chemistry techniques to determine which reactions are most important, rather than waiting for laboratory measurements to be done or using highly inaccurate estimates of the rate of unstudied reactions. This is a new and very much needed approach for studying the atmosphere of Venus," Lyons said. "People are reluctant to go in the lab to measure rate constants for molecules made up of S, chlorine (Cl), and oxygen (O)—these are difficult and sometimes dangerous compounds to work with. Computational methods are the best—and really only—alternative.

Computational methods were used to compute the rate constants and to determine the expected reaction products. These are state-of-the-art computational models (what we call ab initio models). These ab initio calculations were done by the authors from Spain and from the University of Pennsylvania.

"This research illustrates another <u>pathway</u> to S_2 and sulfur particle formation. Sulfur chemistry is dominant in Venus' atmosphere, and very likely plays a key role in the formation of the enigmatic UV absorber. More generally, this work opens the doors to using molecular ab initio techniques to disentangle the complex chemistry of Venus," Lyons said.

More information: Antonio Francés-Monerris et al, Photochemical and thermochemical pathways to S2 and polysulfur formation in the atmosphere of Venus, *Nature Communications* (2022). <u>DOI:</u> <u>10.1038/s41467-022-32170-x</u>

Provided by Planetary Science Institute



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