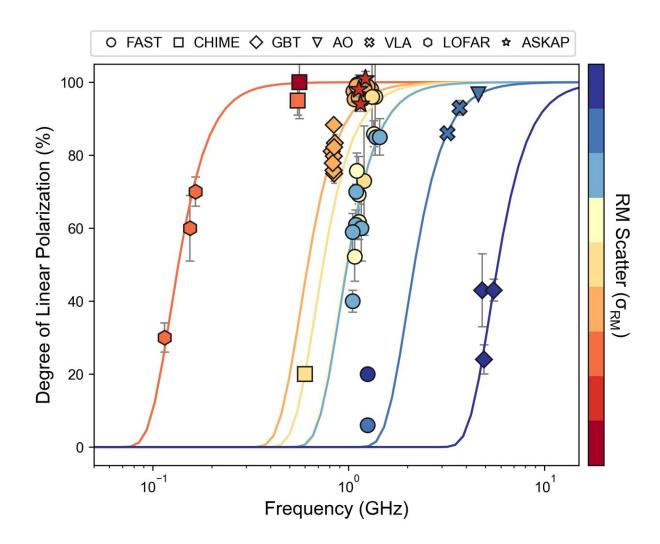


Frequency-dependent polarization of repeating fast radio bursts reveals their origin

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The degree of linear polarization for FRB sources is consistent with RM scattering. Credit: NAOC



Fast radio bursts (FRBs) are the brightest millisecond-duration astronomical transients in radio bands with a yet unknown origin.

The polarization of FRBs contains crucial information about their environments. High-fidelity polarization measurements of FRBs are thus important to understanding their origin.

A research team led by Dr. Li Di from the National Astronomical Observatories of the Chinese Academy of Sciences (NAOC) has analyzed the polarization properties of five repeating FRB sources using the Five-hundred-meter Aperture Spherical radio Telescope (FAST) and the Robert C. Byrd Green Bank Telescope (GBT).

The study was published in *Science* on March 18.

Combined with a diverse set of existing data, these observations reveal systematic frequency evolution, namely depolarization toward <u>lower frequencies</u>, which can be well described by a single parameter, i.e., rotation measure (RM) scattering (σ_{RM}).

Such a unified description of repeating FRBs indicates a complex environment near and/or around the bursting sources, which could be a supernova remnant, a pulsar wind nebula, or plasma near massive blackholes.

FRB 121102, the first repeating FRB source, possesses a high degree of polarization with a correspondingly extremely large RM, which is a product of electron density and magnetic field strength along the line of sight.

According to a study published in *Nature* in 2021, in 50-day span in



2019, FAST detected 1,652 pulses, but no polarization has been detected within this treasure trove. "Non-detection of linear polarization is a norm rather than exception with FAST in its band between 1.0–1.4 GHz, despite its unparalleled sensitivity," said Dr. Li Di, the corresponding author of both papers.

The team has also continued to monitor other active repeaters with FAST. "We were much puzzled by the lack of polarization in most of our observations. Later, when we systematically looked into these FRBs with other major facilities in other <u>frequency bands</u>, particularly those higher than that of FAST, a unified picture emerged," said the first author of the paper, Dr. Feng Yi, currently a staff scientist at the Zhejiang National Lab.

In a combined analysis with other published data amounting to a total of nine repeating and 12 single-burst FRBs, the researchers were able to quantitatively describe the depolarization toward low frequencies based on a simple physical picture, namely multi-path RM scattering. Due to the inhomogeneous electron density in front of and/or around the FRBs, the pulse signal arriving at any given instant could contain radio photons that have traveled along different paths. As a result, the pulse signal loses part/all of its polarization as the photons are aggregated by the telescope receiver.

Such an effect is greater on photons/light waves of longer wavelength (lower frequency) and gives rise to the apparent discrepancies among telescopes. In particular, it explains why the <u>polarization</u> was not detected at FAST.

Such a simple explanation, with only one free parameter, represents a major step toward a physical understanding of the origin of repeating FRBs. The numerical value of RM scatter varies among repeaters, with those with larger σ_{RM} tending to be more active and with signs of a more



complex environment. Yang Yuanpei from Yunnan University has been leading the team's model development efforts. A similar trend may be found in non-repeaters, although such multiple-band information for non-repeaters has not been achieved due to technical difficulties.

"These extremely active FRBs could be a distinct population. We start to see the evolutionary trend in FRBs, with more active sources in more complex environments represented by larger σ_{RM} being younger explosions," said Dr. Li.

More information: Yi Feng et al, Frequency Dependent Polarization of Repeating Fast Radio Bursts - Implications for Their Origin, *Science* (2022). DOI: 10.1126/science.abl7759. www.science.org/doi/10.1126/science.abl7759

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