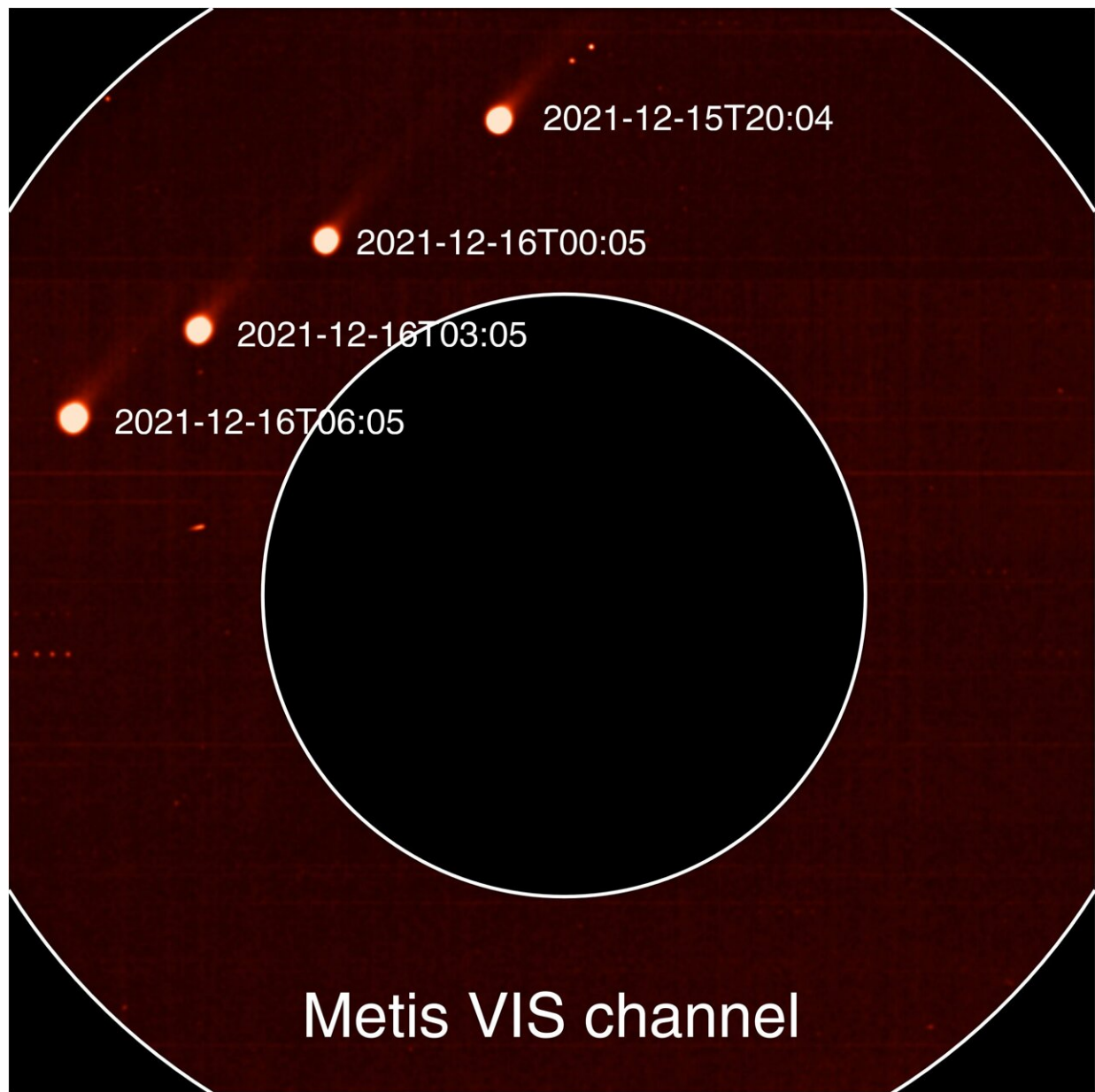


Solar Orbiter catches a second comet by the tail

January 26 2022



A composite of Comet Leonard images captured 15-16 December in visible light by the Metis instrument onboard the ESA/NASA Solar Orbiter spacecraft. The comet transited across the field of view with its dust and ion tails pointing towards the instrument. Credit: ESA/Solar Orbiter/Metis Team

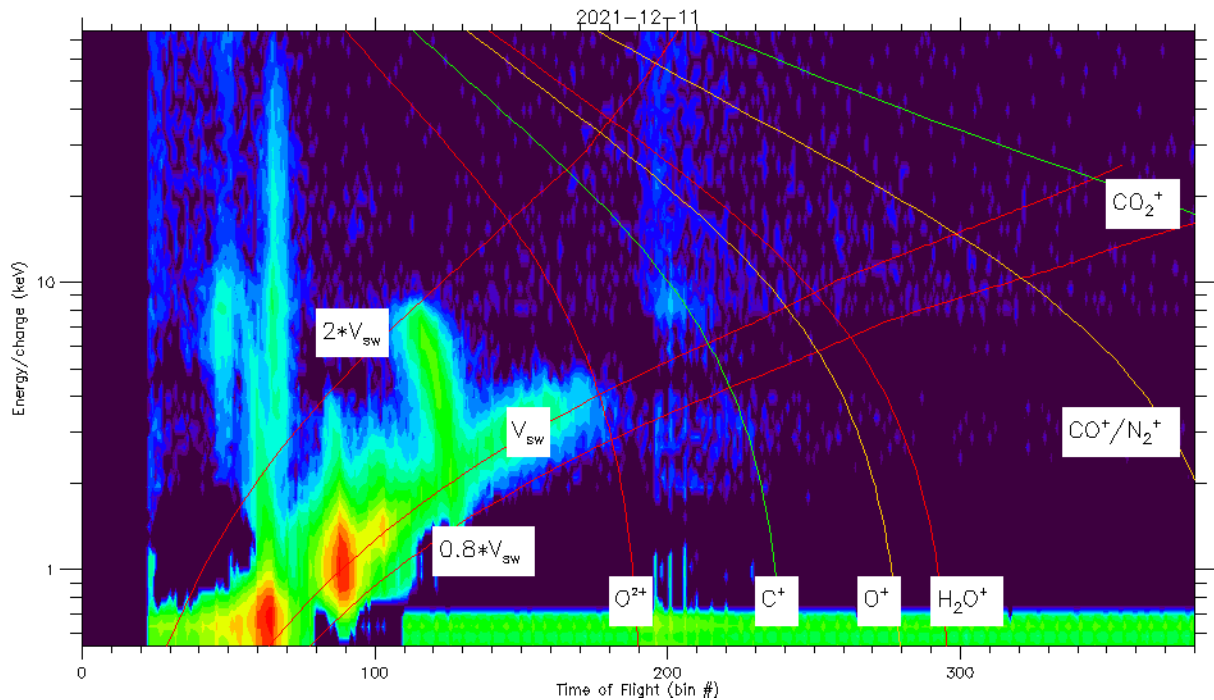
For the second time in its mission so far, the ESA/NASA Solar Orbiter spacecraft has flown through the tail of a comet. Predicted in advance by astronomers at University College London, UK, the spacecraft collected a wealth of science data that now awaits full analysis.

For a spacecraft designed to conduct unique studies of the sun, Solar Orbiter is also making a name for itself exploring comets. For several days centered on 1200-1300 UT on 17 December 2021, the spacecraft found itself flying through the tail of Comet C/2021 A1 Leonard.

The encounter captured information about the particles and [magnetic field](#) present in the tail of the comet. This will allow astronomers to study the way the comet interacts with the [solar wind](#), a variable wind of particles and magnetic field that emanate from the sun and sweep through the solar system.

The crossing had been predicted by Samuel Grant, a post graduate student at University College London's Mullard Space Science Laboratory. He adapted an existing computer program that compared spacecraft orbits with comet orbits to include the effects of the solar wind and its ability to shape a comet's tail.

"I ran it with Comet Leonard and Solar Orbiter with a few guesses for the speed of the solar wind. And that's when I saw that even for quite a wide range of solar wind speeds it seemed like there would be a crossing," he says.



This plot series represents data collected by the Solar Wind Analyser’s Heavy Ion Sensor as the ESA/NASA Solar Orbiter spacecraft passed through Comet Leonard’s tail in December 2021. The data cover 11-20 December, with the first and last plot before and after the tail crossing, respectively, marked by an absence of singly ionized ions. During the tail crossing the instrument detected particles that are attributable to the comet rather than the solar wind, for example ions of oxygen, carbon and molecular nitrogen, and molecules of carbon monoxide, carbon dioxide and water. (Ions are atoms or molecules that have been stripped of one or more electron and now carry a net positive electrical charge.) The plot illustrates the accumulation of ‘hits’ of different cometary and solar particles over the course of each day, and records the time it takes for individual particles to travel through the core of the instrument (time of flight), along with its energy/charge. In general, heavier particles take longer to travel than lighter ones. Particles of different mass/charge distribute along specific lines in the plot. The regular solar wind particles cluster along one line (labelled V_{sw}) in high numbers (indicated by the red and green colours). Boundary lines are also marked for particles travelling twice as fast as solar wind particles ($2V_{sw}$) and 0.8 times as fast ($0.8V_{sw}$). The density of green colours in distinct bands in the right hand side of the plot reveals particles that have different

charge characteristics to those expected in the solar wind, and, in addition, molecules simply not found in the solar wind. Data like this collected over many days helps profile the comet's ion tail. The data are used to determine the rate at which the solar wind is stripping material from the comet, for example by comparing the abundances of 'parent' molecules (e.g. carbon monoxide, CO), with the individual atomic components (in that case, carbon and oxygen). Combined with magnetic field data, scientists will also be able to explore local space plasma instabilities generated by the interaction of the newly created ions with the solar wind itself. Credit: ESA/Solar Orbiter/SWA team

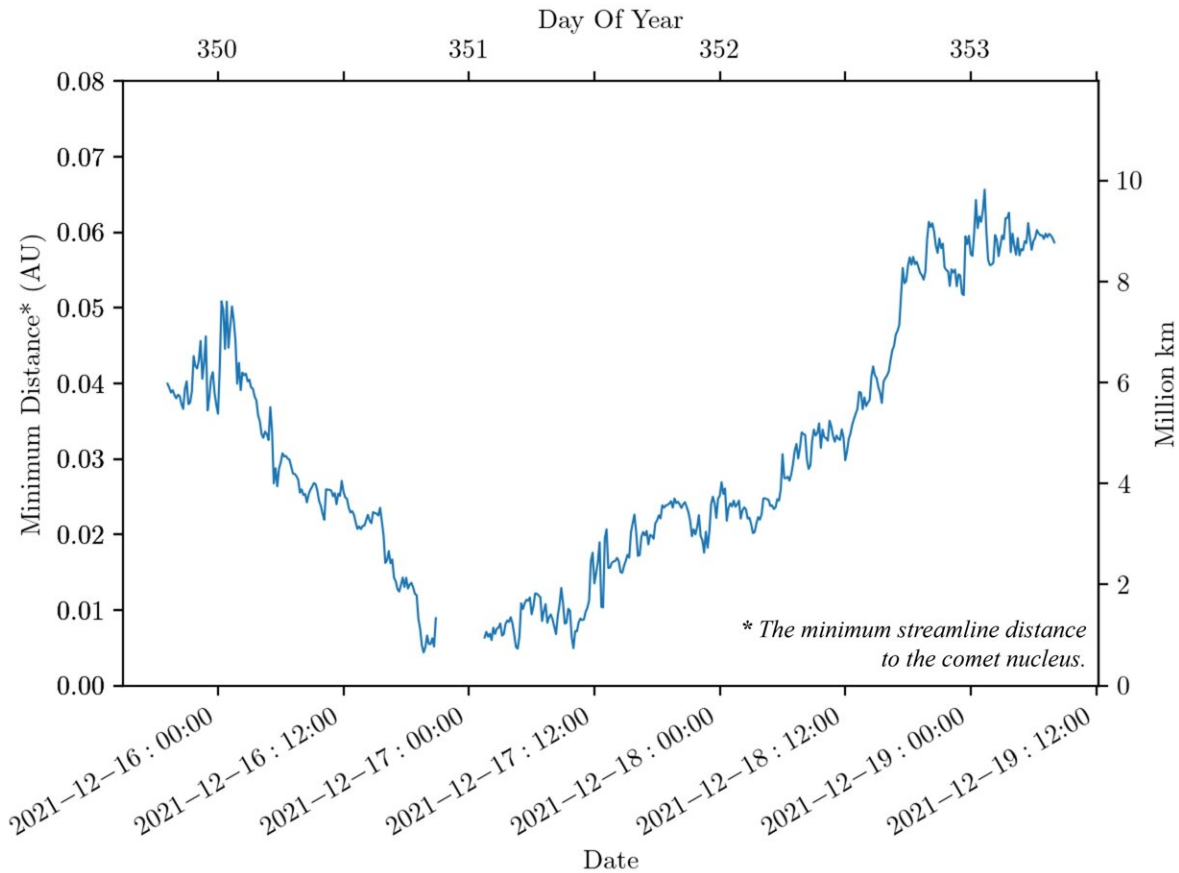
At the time of the crossing, Solar Orbiter was relatively close to the Earth having passed by on 27 November 2021 for a gravity assist maneuver that marked the beginning of the mission's science phase, and placed the spacecraft on course for its March 2022 [close approach](#) to the sun. The comet's nucleus was 44.5 million kilometers away, near to the planet Venus, but its giant tail stretched across space to Earth's [orbit](#) and beyond.

So far, the best detection of the comet's tail from Solar Orbiter has come from the Solar Wind Analyser (SWA) instrument suite. Its Heavy Ion Sensor (HIS) clearly measured atoms, ions and even molecules that are attributable to the comet rather than the solar wind.

Ions are atoms or molecules that have been stripped of one or more electron and now carry a net positive electrical charge. SWA-HIS detected ions of oxygen, carbon, molecular nitrogen, and molecules of carbon monoxide, carbon dioxide and possibly water. "Because of their small charge, these ions are all clearly of cometary origin," says Stefano Livi, Lead Investigator of SWA-HIS from Southwest Research Institute, Texas.

As a comet moves through space, it tends to drape the sun's magnetic

field around it. This magnetic field is being carried by the solar wind, and the draping creates discontinuities where the polarity of the magnetic field changes sharply from north to south and vice versa.



This data plot uses solar wind speed and direction data from the Solar Wind Analyser’s proton and alpha sensor (SWA-PAS) to estimate how close the ESA/NASA Solar Orbiter spacecraft approached to the centre of Comet C/2021 A1 Leonard’s ion tail during December 2021. The plot records how close each packet of solar wind detected by SWA-PAS is thought to have got to the comet’s nucleus during its journey from the Sun to the spacecraft. The left axis gives the scale in astronomical units (au), where 1 au is the distance from the Sun to the Earth, and the same distance is shown in kilometres on the right axis. Changes in the solar wind flow speed and direction are responsible for the variations in the

plotted distance. There are short data acquisition gaps on 15 and 17 December. Data like these can help determine the timing of the tail crossing. Comet ion detections from another SWA sensor (the Heavy Ion Sensor) lasted several days, demonstrating that the tail was millions of kilometres across. Homing in on the mid-point of the tail-crossing based on the first analysis puts it from the end of 16 December (according to the data plot shown here) to 1200-1300 UT on 17 December (based on the preliminary analysis of signatures from other instruments). Credit: ESA/Solar Orbiter/SWA team & S. Grant (UCL)

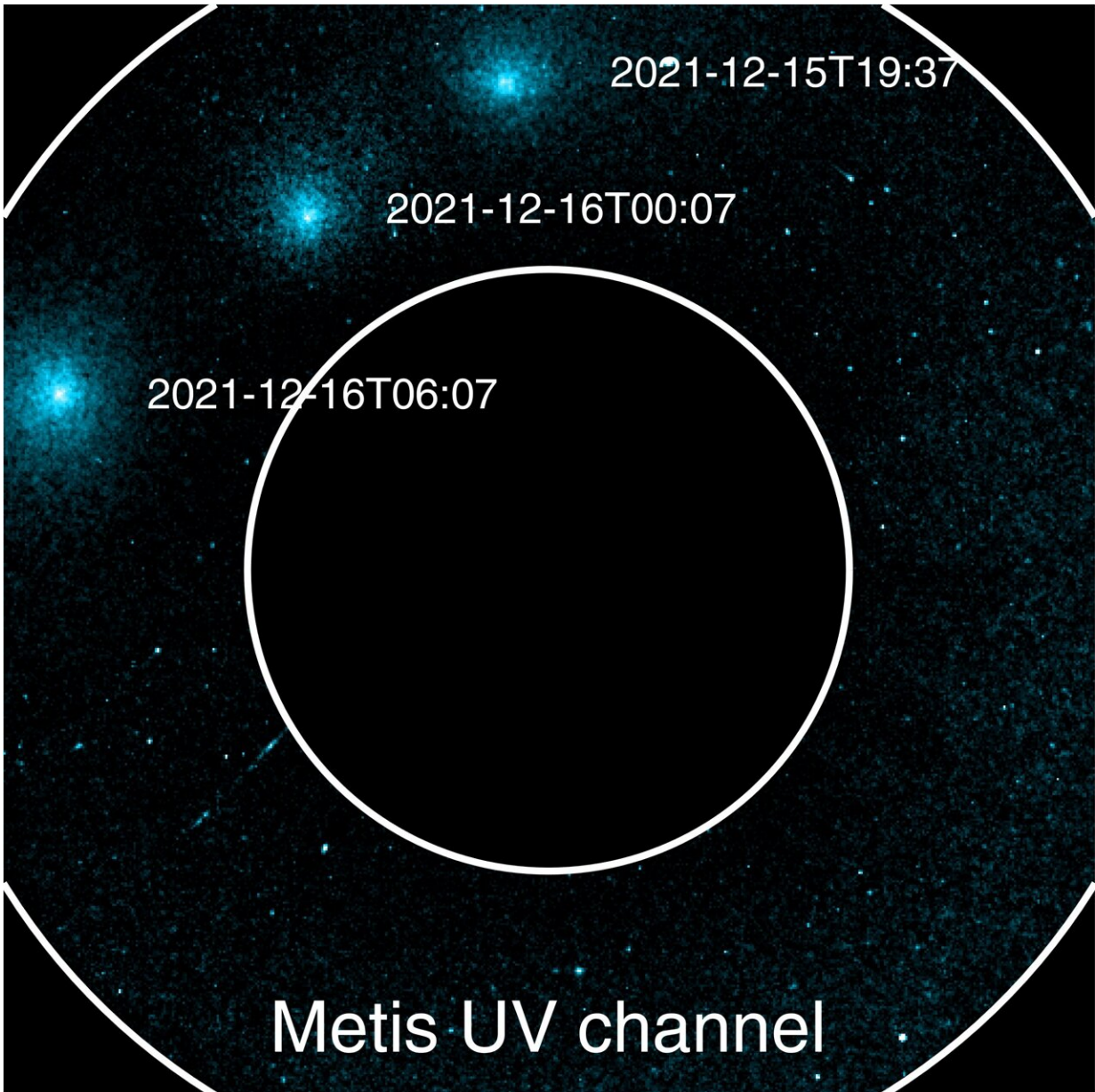
The magnetometer instrument (MAG) data does indeed suggest the presence of such draped magnetic field structures but there is more analysis to be done to be absolutely sure. "We are in the process of investigating some smaller scale magnetic perturbations seen in our data and combining them with measurements from Solar Orbiter's particle sensors to understand their possible cometary origin," says Lorenzo Matteini, a co-investigator on MAG from Imperial College, London.

In addition to the particle data, Solar Orbiter also acquired images.

Metis is Solar Orbiter's multi-wavelength coronagraph. It can perform ultraviolet observations that see the Lyman alpha emission given out by hydrogen, and it can measure the polarization of visible light. During 15 and 16 December it captured the distant head of the comet simultaneously in both visible and ultraviolet light. These images are now being analyzed by the instrument team. "The visible light images can hint at the rate at which the comet is ejecting dust, while the ultraviolet images can give the water production rate," says Alain Corso, a Metis co-investigator at the CNR-Istituto di Fotonica e Nanotecnologie, Padova, Italy.

The Solar Orbiter Heliospheric Imager (SoloHI) also captured data. These images show large parts of the comet's ion tail taken while the

spacecraft itself was inside the tail. As the image sequence progresses, changes in the tail can be seen in response to variations in the solar wind speed and direction.



A composite of Comet Leonard images captured 15-16 December in ultraviolet light by the Metis instrument onboard the ESA/NASA Solar Orbiter spacecraft. The comet transited across the field of view with its dust and ion tails pointing

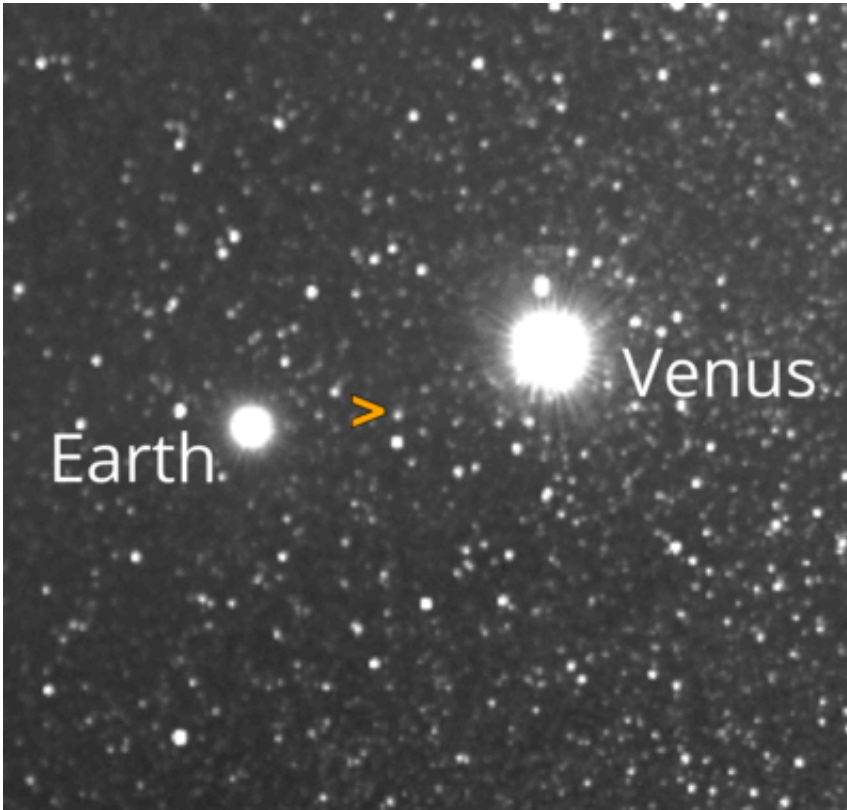
towards the instrument. Credit: ESA/Solar Orbiter/Metis Team

And it was not just Solar Orbiter that was watching the crossing. The ESA/NASA SOHO mission and NASA's STEREO-A and Parker Solar Probe spacecraft were observing from afar. This means that not only do astronomers now have data from inside the tail, they also have contextual images from these other spacecraft (see image gallery above).

Comet tail crossings are relatively rare events. Of those that have been detected, most have been noticed only after the event. The ESA/NASA Ulysses mission encountered three comet ion tails, including that of C/1996 B2 Hyakutake in May 1996, and C/2006 P1 McNaught in early 2007. Solar Orbiter itself crossed the tail of fragmenting comet C/2019 Y4 ATLAS in May and June 2020, shortly after launching.

Whereas the early crossings were a surprise, both of Solar Orbiter's encounters were predicted in advance thanks to the computer code developed by Geraint Jones, University College London Mullard Space Science Laboratory, and extended by Samuel.

"The big advantage is that for basically no effort on the spacecraft's part, you get to sample a comet at a massive distance. That's pretty exciting," says Samuel, who is now looking at archive data from other spacecraft looking for comet [tail](#) crossings that have so far gone unnoticed.

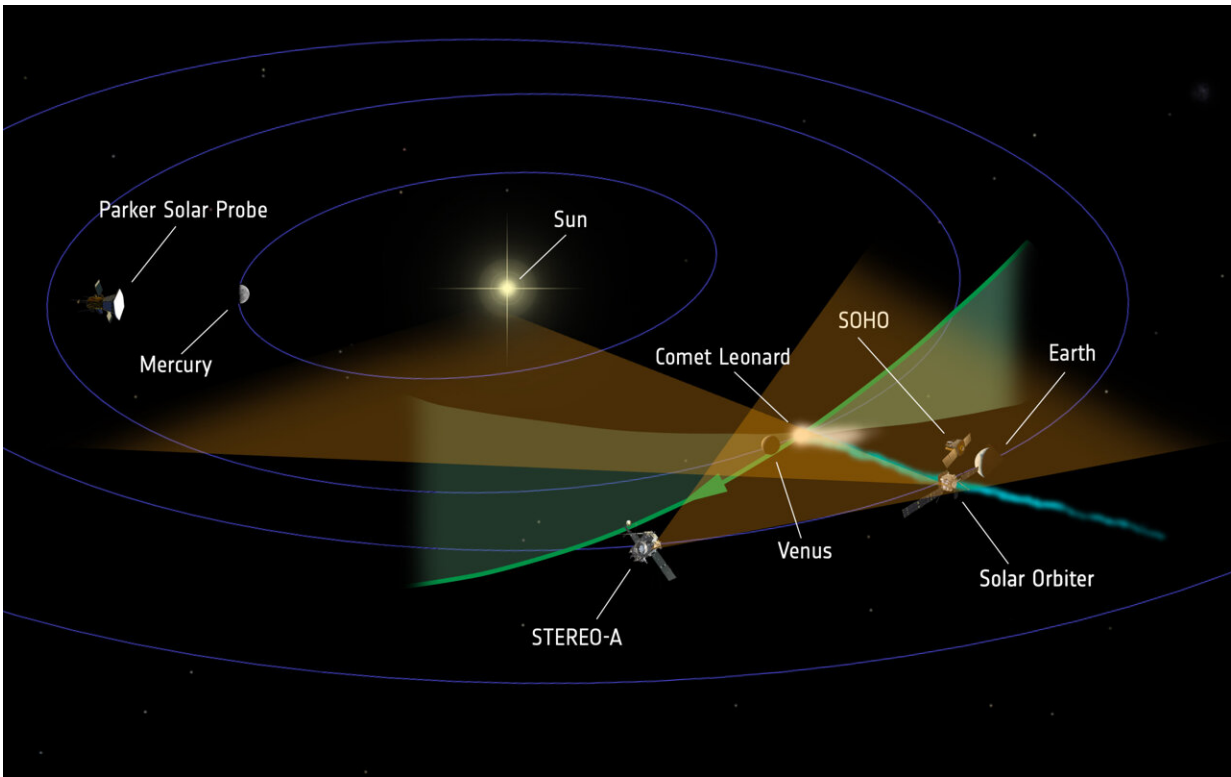


NASA's Parker Solar Probe was almost on the opposite side of the Sun from the ESA/NASA Solar Orbiter spacecraft when its WIPSR instrument recorded images of Comet Leonard on 7 December 2021. From Parker Solar Probe's viewpoint the comet appeared to pass directly between Venus and Earth. Credit: NASA/USNRL/G.Stenborg/K.Battams

The work also helps build experience for ESA's Comet Interceptor mission, for which Geraint is the Science Team Lead. The mission will visit an as-yet undiscovered comet, making a flyby of the target with three spacecraft to create a 3D profile of a 'dynamically new' object that contains unprocessed material surviving from the dawn of the [solar system](#).

In the meantime, the instrument teams on Solar Orbiter are busy analyzing the Comet Leonard data not only for what it can tell them

about the comet but about the solar wind as well.



The ESA/NASA Solar Orbiter spacecraft flew through the tail of Comet C/2021 A1 Leonard in December 2021, collecting images and in-situ solar wind and particle data. At the same time, SOHO (ESA/NASA), Parker Solar Probe (NASA) and STEREO-A (NASA) were also watching the comet's evolution from other angles. The graphic shows the approximate relative positions of the planets, comet and spacecraft on 17 December 2021 and is not to scale. Very approximate fields of view are indicated for selected instruments: SoloHI on Solar Orbiter and SECCHI on STEREO-A. Credit: G. Jones & S. Grant (UCL)

"This kind of additional science is always an exciting part of a space mission," says Daniel Müller, ESA Project Scientist for Solar Orbiter. "When the comet ATLAS crossing was predicted, we were still

calibrating the spacecraft and its instruments. Also, the comet fragmented just before we got there. But with Comet Leonard we were totally ready—and the [comet](#) didn't fall apart."

In March, Solar Orbiter make its closest pass to the sun yet at a distance of 0.32 au (approximately one-third of the Earth-sun distance, or about 50 million kilometers). It is one of almost 20 close passes to the sun that will occur during the next decade. These will result in unprecedented images and data, not only from close up, but also from the sun's never-before seen polar regions.

"There is so much to look forward to with Solar Orbiter, we're only just getting started," says Daniel.

Provided by European Space Agency

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