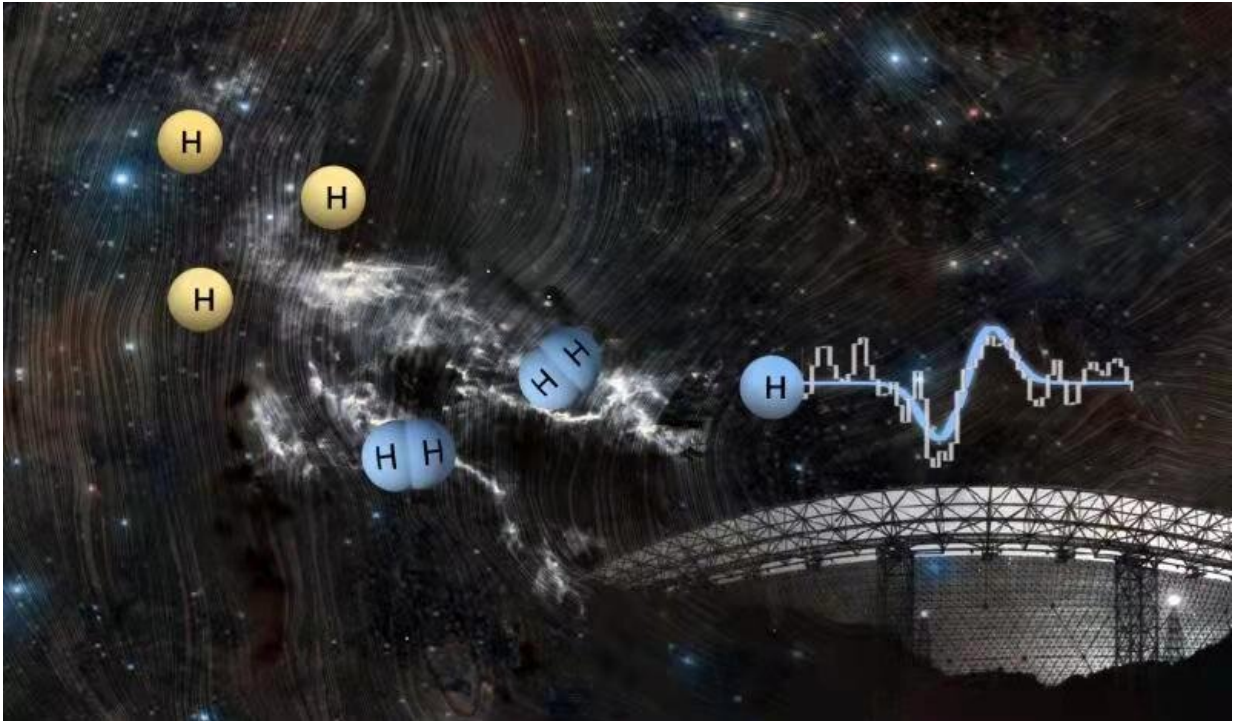


# Coherent interstellar magnetic field detected

January 5 2022

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The Taurus molecular cloud (grey scale), of which L1544 is a part, is superimposed onto the 2MASS sky image and the field orientation based on Planck data (thin white lines). The HINSA Zeeman spectrum (thick white line) is shown with the fitted Zeeman signature (blue). Credit: NAOC

Magnetic fields are the essential, but often "secret" ingredients of the interstellar medium and the process of making stars. The secrecy shrouding interstellar magnetic fields can be attributed to the lack of experimental probes.

While Michael Faraday was probing the link between magnetism and electricity with coils in the early 19th century in the basement of the Royal Institution, astronomers today still cannot deploy coils light-years away.

Using the Five-hundred-meter Aperture Spherical radio Telescope (FAST), an international team led by Dr. LI Di from National Astronomical Observatories of Chinese Academy of Sciences (NAOC) has obtained accurate [magnetic field](#) strength in molecular cloud L1544—a region of the interstellar medium that seems ready to form stars.

The team employed the so-called HI Narrow Self-Absorption (HINSA) technique, first conceived by LI Di and Paul Goldsmith based on Arecibo data in 2003. FAST's sensitivity facilitated a clear detection of the HINSA's Zeeman effect. The results suggest that such clouds achieve a supercritical state, i.e., are primed for collapse, earlier than standard models suggest.

"FAST's design of focusing [radio waves](#) on a cable-driven cabin results in clean optics, which has been vital to the success of the HINSA Zeeman experiment," said Dr. LI.

The study was published in *Nature* on Jan. 5.

The Zeeman effect—the splitting of a spectral line into several components of frequency in the presence of a magnetic field—is the only direct probe of interstellar magnetic field strength. The interstellar Zeeman effect is small. The frequency shift originating in the relevant clouds is only a few billionths of the intrinsic frequencies of the emitting lines.

In 2003, the spectra of molecular clouds were found to contain an

atomic-hydrogen feature called HINSA, which is produced by hydrogen atoms cooled through collisions with hydrogen molecules. Since this detection was made by the Arecibo telescope, the Zeeman effect for HINSA has been deemed a promising probe of the magnetic field in molecular [clouds](#).

HINSA has a line strength 5–10 times higher than that of molecular tracers. HINSA also has a relatively strong response to magnetic fields and, unlike most molecular tracers, is robust against astrochemical variations.

FAST's HINSA measurements put the magnetic field strength in L1544 at about 4  $\mu$ Gauss, i.e., 6 million times weaker than that of Earth. A combined analysis with quasar (active supermassive blackhole) absorption and hydroxyl emission also revealed a coherent magnetic field structure throughout the cold neutral medium, the molecular envelope, and the dense core, with similar orientation and magnitude.

Therefore, the transition from magnetic subcriticality to supercriticality—i.e., when the field can and cannot support the cloud against gravity, respectively—occurs in the envelope instead of the core, in contrast with the conventional picture.

How the interstellar magnetic field dissipates to enable cloud collapse remains an unsolved problem in star formation. The main proposed solution has long been ambipolar diffusion—the decoupling of neutral particles from plasma—in cloud cores.

The coherence of the magnetic field revealed by the HINSA Zeeman effect means that dissipation of the [field](#) occurs during the formation of the molecular envelope, possibly through a different mechanism than ambipolar diffusion.

**More information:** Tao Ching, An early transition to magnetic supercriticality in star formation, *Nature* (2022). [DOI: 10.1038/s41586-021-04159-x](https://doi.org/10.1038/s41586-021-04159-x).  
[www.nature.com/articles/s41586-021-04159-x](https://www.nature.com/articles/s41586-021-04159-x)

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