

Satellite galaxies can continue forming stars when they pass close to their parent galaxies

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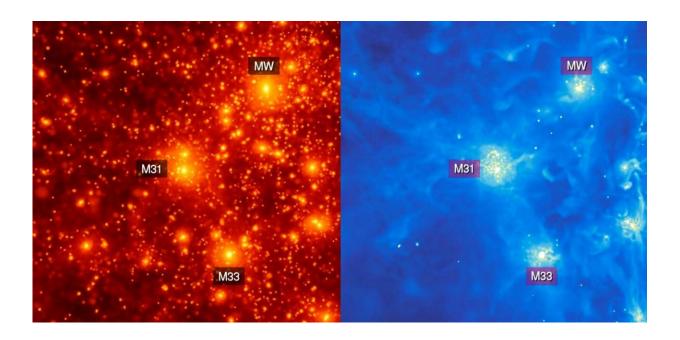


Image of the simulated local group used for the article. Left, image of dark matter; on the right, gas distribution. The three main galaxies of the Local Group (MW, M31 and M33) are indicated. Credit: CLUES simulation team

Historically most scientists thought that once a satellite galaxy has passed close by its higher mass parent galaxy, its star formation would stop because the larger galaxy would remove the gas from it, leaving it shorn of the material it would need to make new stars. However, for the first time, a team led by Arianna di Cintio, a researcher at the Instituto de Astrofísica de Canarias (IAC), has shown using numerical simulations



that this is not always the case. The results of the study were recently published in the journal *Monthly Notices of the Royal Astronomical Society (MNRAS)*.

Using sophisticated simulations of the whole of the Local Group of galaxies, including the Milky Way, the Andromeda galaxy and their respective satellite galaxies, the researchers have shown that the satellites not only can retain their gas but can also experience many new episodes of star formation just after passing close to the pericenter of their parent galaxy (the mínimum distance they reach from its center).

The satellite galaxies of the Local Group show a wide variety of star formation histories, whose origin has not previously been fully understood. Using hydrodynamic simulations within the project Constrained Local UniversE (CLUES) the authors studied the star formation histories of satellite galaxies similar to those of the Milky Way in a cosmological context.

While in the majority of the cases the gas of the satellite is sucked out by the parent galaxy due to gravitational action and transfers itself to the larger galaxy (interrupting star formation in the satellite, in a process known as accretion), in some 25% of the sample they found that star formation was clearly enhanced by this interactive process.

The results show that the peaks of star formation are correlated with the close pass of the satellite around the parent galaxy, and occasionally by the interaction of two satellites. The researchers identified two key features to the star formation: The satellite must enter the parent galaxy with a large reserve of cold gas, and a minimum distance not too small, so that <u>stars</u> may form due to compression of the gas. On the contrary, galaxies which pass too close to the parent galaxy, or to a parent galaxy with little gas, are stripped of their gas and thereby lose the possibility of forming <u>new stars</u>.



"The passage of satellites also coincide with peaks in the star formation of their parent galaxies, which suggests that this mechanism causes bursts of stars equally in both parent galaxies and satellites, in agreement with recent studies of the history of star formation in our own galaxy," explains Arianna di Cintio, the lead author on the paper.

"This is very important when we try to understand how star formation is produced in the smaller dwarf galaxies of our Local Group, an unresolved question," she adds.

This finding will shed light on the episodes of star formation which are observed in the dwarf galaxies of the Local Group, such as Carina and Fornax, giving an attractive explanation of their existence. It also requires a revision of the theoretical models used to explain the formation of stars in dwarf galaxies.

More information: Arianna Di Cintio et al, Pericentric passage-driven star formation in satellite galaxies and their hosts: CLUES from Local Group simulations, *Monthly Notices of the Royal Astronomical Society* (2021). DOI: 10.1093/mnras/stab1682

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