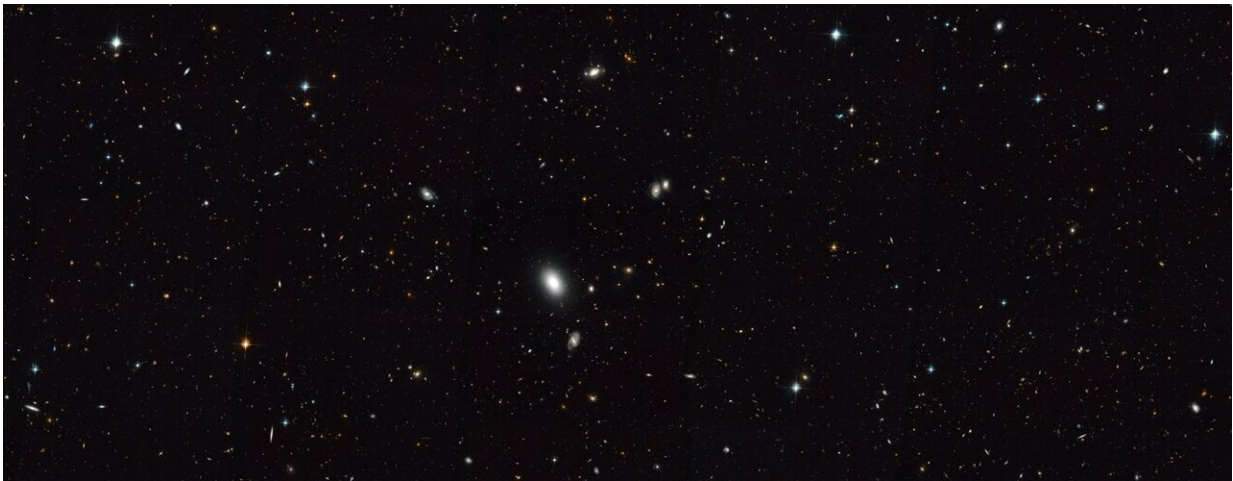


Mapping the early universe with NASA's Webb Telescope

June 25 2020, by Claire Blome



The CEERS Survey researchers will use the James Webb Space Telescope to observe the Extended Groth Strip in infrared light. Their observations employ three of the telescope's instruments and will provide both images and spectra of the objects in the field — which includes at least 50,000 galaxies — helping to expand what we know about galaxies in the very early universe. Credit: NASA, ESA, and M. Davis (University of California, Berkeley)

Astronomers and engineers have designed telescopes, in part, to be "time travelers." The farther away an object is, the longer its light takes to reach Earth. Peering back in time is one reason why NASA's upcoming James Webb Space Telescope specializes in collecting infrared light: These longer wavelengths, which were initially emitted by stars and

galaxies as ultraviolet light more than 13 billion years ago, have stretched, or redshifted, into infrared light as they traveled toward us through the expanding universe.

Although many other observatories, including NASA's Hubble Space Telescope, have previously created "deep fields" by staring at small areas of the sky for significant chunks of time, the Cosmic Evolution Early Release Science (CEERS) Survey, led by Steven L. Finkelstein of the University of Texas at Austin, will be the first for Webb. He and his research team will spend just over 60 hours pointing the telescope at a slice of the sky known as the Extended Groth Strip, which was observed as part of Hubble's Cosmic Assembly Near-infrared Deep Extragalactic Legacy Survey or CANDELS.

"With Webb, we want to do the first reconnaissance for [galaxies](#) even closer to the big bang," Finkelstein said. "It is absolutely not possible to do this research with any other telescope. Webb is able to do remarkable things at wavelengths that have been difficult to observe in the past, on the ground or in space."

Mark Dickinson of the National Science Foundation's National Optical-Infrared Astronomy Research Laboratory in Arizona, and one of the CEERS Survey co-investigators, gives a nod to Hubble while also looking forward to Webb's observations. "Surveys like the Hubble Deep Field have allowed us to map the history of cosmic star formation in galaxies within a half a billion years of the big bang all the way to the present in surprising detail," he said. "With CEERS, Webb will look even farther to add new data to those surveys."

Delivering the Unseen

What was the [early universe](#) like? There are certainly many [data points](#), but not enough to create an exhaustive census of its conditions. Plus,

researchers' knowledge and assumptions are updated frequently—each time a new deep exposure is released. "Every time we look farther, we find galaxies earlier and earlier than we thought possible. The conditions in the very early [universe](#) had to be right for galaxies to form—and they formed and became massive very quickly," said CEERS Survey co-investigator Jeyhan Kartaltepe of the Rochester Institute of Technology in New York.

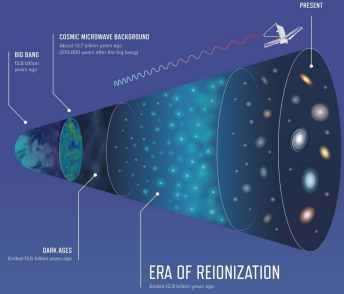
"The universe was more compact at this time, which means stars and galaxies could have formed at a greater efficiency," Finkelstein added. "Some models predict we'll find 50 galaxies at earlier eras more distant than Hubble can reach, but others predict we will only find a few. In both cases, the data will help us constrain galaxy formation in the early universe."

The CEERS Survey team hopes to identify an abundance of distant objects, including the most distant galaxies in the universe, early galaxy mergers and interactions, the first massive or supermassive black holes, and even earlier quasars than previously identified. These potential "firsts" are only the beginning of the value of this research: The team, which is made up of over 100 researchers from around the world, will go on to classify many objects in the field. "These data will help demonstrate what the structure of the universe was like at various periods," Finkelstein explained.

Cosmic Reionization

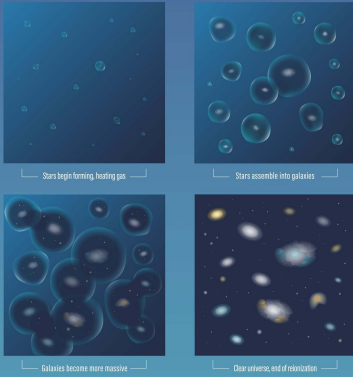
EXPLORING THE EARLY UNIVERSE

Our view of the universe wasn't always so clear. More than 13 billion years ago, neutral gas made the universe opaque to some types of light. Over hundreds of millions of years, the universe became transparent as its gas particles became charged or ionized. What caused the gas to change? The James Webb Space Telescope will peer deep into space to gather more information about objects in this period, known as the Era of Reionization, to help us understand the major transition in the history of the universe.



WHAT WE DO KNOW

After the first stars formed, the universe was still lost in a gaseous fog, but as stars and young galaxies continued to evolve and produce more energetic light, they began to change the gas around them—converting it from neutral to ionized gas. Eventually, they transformed the space, making it possible to observe these early galaxies.



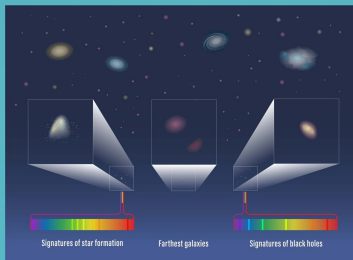
OODLES OF NEW DATA

How did stars and galaxies change over time? Webb's infrared observations will help us create the first detailed snapshot of galaxies in the early universe—as well as provide much more information than ever possible before. With this new data, researchers will begin to analyze individual objects to understand how the surrounding gas changed from neutral to ionized, creating the transparent universe we see today.



LOOKING BACK IN TIME

Webb will dig deeper into the universe's history than any other telescope—showing us galaxies as they began assembling more than 13 billion years ago. Only Webb has the sensitivity and resolution to detect higher radiation energies and unique, colorful spectra in infrared light (depicted below as rainbows) that will help us learn about the conditions in the Era of Reionization.



More than 13 billion years ago, during the Era of Reionization, the universe was a very different place. The gas between galaxies was largely opaque to energetic light, making it difficult to observe young galaxies. What allowed the universe to become completely ionized, or transparent, eventually leading to the “clear” conditions detected in much of the universe today? The James Webb Space Telescope will peer deep into space to gather more information about objects that existed during the Era of Reionization to help us understand this major transition in the history of the universe. Credit: NASA, ESA, and J. Kang (STScI)

Hitting "Rewind"

Perhaps the most exciting element of this research is how the team will use the data to uncover new findings about an important period of the universe's history called the "Era of Reionization." The [big bang](#) set off a series of events, leading to the cosmic microwave background, the dark ages, the first stars and galaxies—and then to the Era of Reionization. During this period, the gas in the universe transformed from mostly neutral, meaning it was opaque to ultraviolet light, and became completely ionized, which allowed it to be transparent. Ionization means the atoms were stripped of their electrons—eventually leading to the "clear" conditions detected in much of the universe today.

Many questions remain about this unique time in our universe. For example, what was responsible for converting the gas from neutral to ionized? And how long did it take before the universe became significantly less opaque and much more transparent?

"We think this happened when ultraviolet light escaped young, forming galaxies," Dickinson explained. "There may be other factors. For

example, early accreting black holes may also have emitted ultraviolet light that eventually helped transform the gas."

Where the galaxies appear on the sky offers another clue. "We'll examine reionization-era galaxies to see if they are clustered together in the same regions or if they are more isolated," said Kartaltepe. "We have a lot of ideas about what causes galaxies to grow and become more massive, but we need more comprehensive information about these galaxies to fully understand how they initially grew and evolved."

The presence of galactic mergers or interactions—or lack thereof—will also help the team trace the conditions of the environment during the Era of Reionization. "The CEERS Survey will give us hints about how this period proceeded," Dickinson adds. "We will certainly learn about the galaxies we think are responsible, and also hope to learn about the ionizing radiation that escaped them."

The team has designed the CEERS Survey to provide as much complementary data as possible for many targets in this field of view. They will employ three of Webb's instruments, in several modes, to obtain images of the Extended Groth Strip, in addition to spectra. Spectra are invaluable data since they help researchers identify the colors, temperatures, motions, and masses of each target, and provide a much more in-depth look at the chemical makeup of distant objects.

"That's the difference with Webb's Near-Infrared Spectrograph, or NIRSpec," Dickinson emphasized. "We'll open the spectrograph's microshutter slits to individually observe hundreds of galaxies to obtain their spectra for the first time."

Beginning to Build a Census

In the months following the initial data release, the CEERS Survey

researchers will create and post new tools and catalogs any researcher can use to analyze the data, including masses of galaxies, galaxy shapes, and photometric redshifts. "With the same set of observations, hundreds of researchers can conduct hundreds of science experiments," Kartaltepe said. "We're also going to find things we didn't even think to ask, which is one more reason why the CEERS Survey research will be so rewarding. Our hope is that the CEERS Survey will influence future distant galaxy surveys with Webb," Finkelstein added. "It will also demonstrate to the community that observing with a variety of instruments and modes are very valid ways to increase Webb's scientific yield."

More information: Cosmic Evolution Early Release Science Survey: www.stsci.edu/jwst/observing-programs/program-1345

Provided by NASA's Goddard Space Flight Center

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