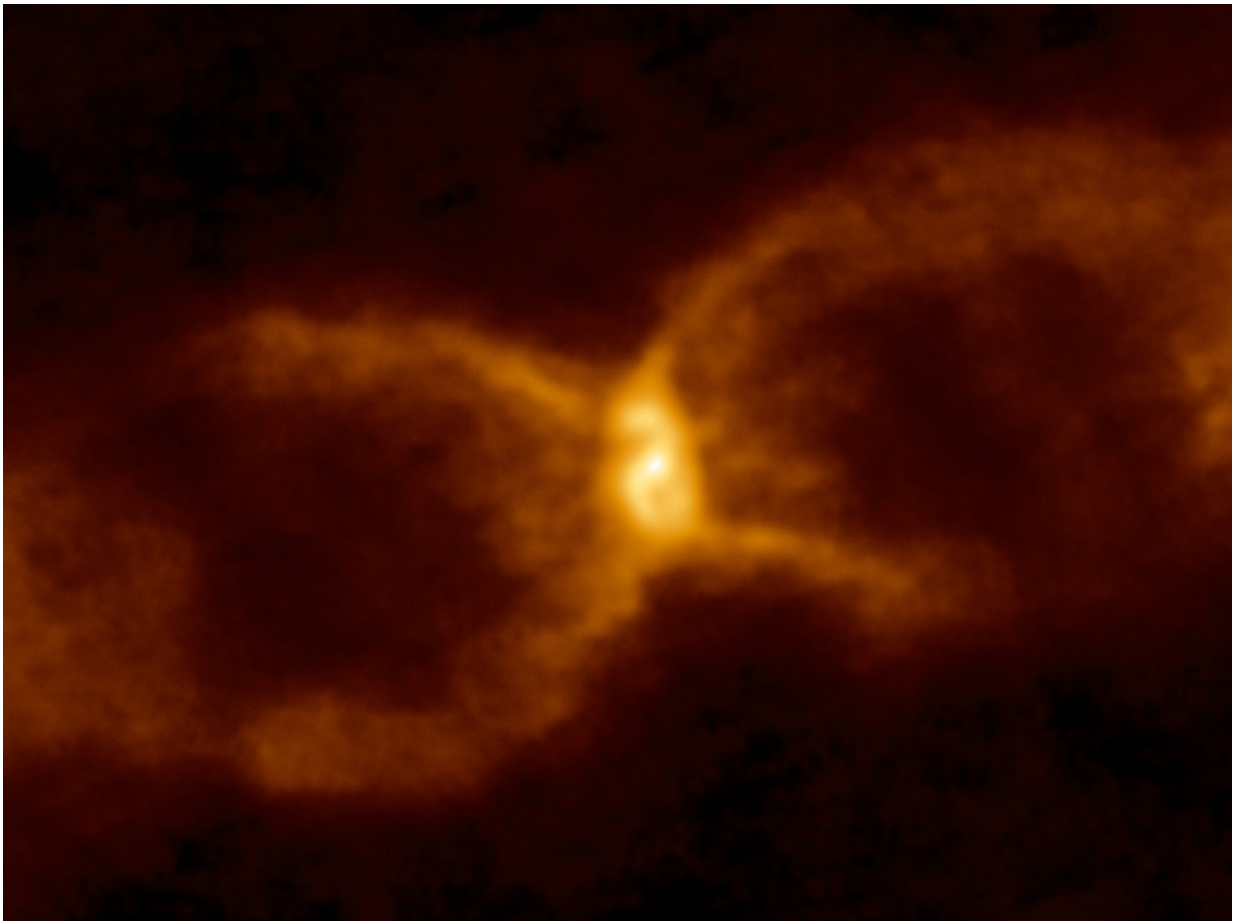


# When is a nova not a nova? When a white dwarf and a brown dwarf collide

October 8 2018

---



This object is possibly the oldest of its kind ever catalogued: the hourglass-shaped remnant named CK Vulpeculae. Credit: ALMA (ESO/NAOJ/NRAO)/S. P. S. Eyres

Researchers from Keele University and an international team of astronomers have reported for the first time that a white dwarf and a brown dwarf collided in a 'blaze of glory' that was witnessed on Earth in 1670.

Using the Atacama Large Millimeter/submillimeter Array (ALMA) in Chile, the astronomers found evidence that a white dwarf (the remains of a sun-like star at the end of its life) and a brown dwarf (a failed star without sufficient mass to sustain thermonuclear fusion) collided in a short-lived blaze of glory that was witnessed on Earth in 1670 as Nova Cygni—"a new star below the head of the Swan." It appeared abruptly as a star as bright as those in the plough, that gradually faded, reappeared, and finally disappeared from view.

Modern astronomers studying the remains of this cosmic event initially thought it was triggered by the merging of two main-sequence [stars](#) on the same evolutionary path as our sun. This nova was long referred to as "Nova Vulpeculae 1670," and later became known as CK Vulpeculae. However, we now know that CK Vulpeculae was not what we would today describe as a nova, but was, in fact, the merger of two stars—a white dwarf and a brown dwarf.

By studying the debris from this explosion, which today appears as dual rings of dust and gas resembling an hourglass with a compact central object, the research team concluded that a brown dwarf merged with a white dwarf. Professor Nye Evans, Professor of Astrophysics at Keele University and co-author on the in the *Monthly Notices of the Royal Astronomical Society*, explains, "CK Vulpeculae has in the past been regarded as the oldest 'old nova.' However, the observations of CK Vulpeculae I have made over the years using telescopes on the ground and in space convinced me that this was no nova. Everyone knew what it wasn't—but nobody knew what it was. But a stellar merger of some sort seemed the best bet. With our ALMA observations of the exquisite dusty

hourglass and the warped disc, plus the presence of lithium and peculiar isotope abundances, the puzzle fit together: In 1670, a brown dwarf star was shredded and dumped on the surface of a [white dwarf star](#), leading to the 1670 eruption and the hourglass we see today."

The team of European, American and South African astronomers used the Atacama Large Millimeter/submillimeter Array to examine the remains of the merger and reported some interesting findings. By studying the light from two more distant stars as they shine through the dusty remains of the merger, the researchers were able to detect the telltale signature of the element lithium, which is easily destroyed in stellar interiors.

Dr. Stewart Eyres, deputy dean of the Faculty of Computing, Engineering and Science at the University of South Wales and lead author on the paper, says, "The material in the hourglass contains the element lithium, normally easily destroyed in stellar interiors. The presence of lithium, together with unusual isotopic ratios of the elements C, N, O, indicate that an astronomically small amount of material, in the form of a [brown dwarf star](#), crashed onto the surface of a white dwarf in 1670, leading to thermonuclear burning, an eruption that led to the brightening seen by the Carthusian monk Anthelme and the astronomer Hevelius, and in the hourglass we see today."

Professor Albert Zijlstra, from The University of Manchester's School of Physics & Astronomy, co-author of the study, says, "Stellar collisions are the most violent events in the universe. Most attention is given to collisions between neutrons stars, or between two [white dwarfs](#)—which can produce a supernova—and star-planet collisions. But it is very rare to actually see a collision, and where we believe one occurred, it is difficult to know what kind of stars collided. The collision here is a new one, not previously considered or ever seen before. This is an extremely exciting discovery."

Professor Sumner Starrfield, Regents' Professor of Astrophysics at Arizona State University says, "The white dwarf would have been about 10 times more massive than the brown dwarf, so as the brown dwarf spiraled into the white dwarf it would have been ripped apart by the intense tidal forces exerted by the white dwarf. When these two objects collided, they spilled out a cocktail of molecules and unusual element isotopes. These [organic molecules](#), which we could detect with ALMA, expanded measurably into the surrounding environment, providing compelling evidence of the true origin of this blast. This is the first time such an event has been conclusively identified. Intriguingly, the hourglass is also rich in organic molecules such as formaldehyde ( $\text{H}_2\text{CO}$ ), methanol ( $\text{CH}_3\text{OH}$ ) and methanamide ( $\text{NH}_2\text{CHO}$ ). These molecules would not survive in an environment undergoing nuclear fusion and must have been produced in the debris from the explosion. This lends further support to the conclusion that a brown dwarf met its demise in a star-on-star collision with a white dwarf."

Since most star systems in the Milky Way are binary, stellar collisions are not that rare, the astronomers note. Professor Starrfield says, "Such collisions are probably not rare, and this material will eventually become part of a new planetary system, implying that they may already contain the building-blocks of organic molecules as they are forming."

**More information:** S P S Eyres et al, ALMA reveals the aftermath of a white dwarf–brown dwarf merger in CK Vulpeculae, *Monthly Notices of the Royal Astronomical Society* (2018). [DOI: 10.1093/mnras/sty2554](https://doi.org/10.1093/mnras/sty2554)

Provided by Keele University

Citation: When is a nova not a nova? When a white dwarf and a brown dwarf collide (2018, October 8) retrieved 20 April 2024 from <https://phys.org/news/2018-10-nova-white-dwarf->

[brown-collide.html](#)

This document is subject to copyright. Apart from any fair dealing for the purpose of private study or research, no part may be reproduced without the written permission. The content is provided for information purposes only.