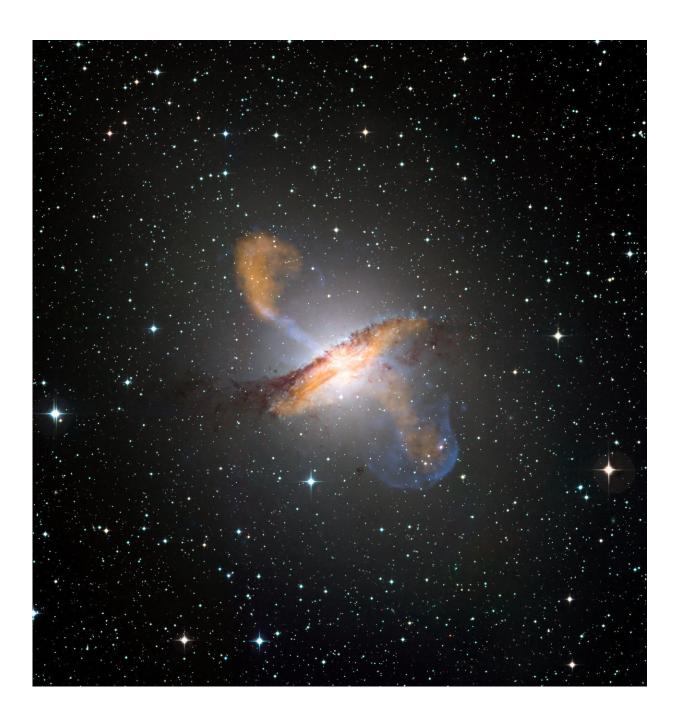


Groundbreaking study sheds new light on galaxy evolution

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Colour composite image of Centaurus A, revealing the lobes and jets emanating from the active galaxy's central black hole. Credit: ESO/WFI (Optical); MPIfR/ESO/APEX/A.Weiss et al. (Submillimetre); NASA/CXC/CfA/R.Kraft et al. (X-ray)

Using integral field spectroscopy (IFS) and advanced modeling tools, Instituto de Astrofísica e Ciências do Espaço (IA) researchers Iris Breda and Polychronis Papaderos have achieved an important milestone towards solving a long standing enigma in extragalactic astronomy – the nature and formation of the central spherical component in spiral galaxies like the Milky Way.

The <u>bulge</u> is thought to form through two distinct routes: Classical bulges consist of ancient stars, older than the disk, because they assembled rapidly more than 10 billion years ago, prior to disks. Pseudo-bulges have stars of similar age as the disk, because they assembled gradually through a combination of dynamical processes, with continuous <u>star</u> formation fed by inflow of gas from the disk.

These two scenarios imply that classical bulges and pseudo-bulges have strikingly different characteristics, but this sharp contrast was never observed, despite numerous studies over the past years.

To solve this riddle, the team has carried out an unprecedented spectral modeling analysis of more than half a million individual spectra, to spatially resolve the star formation history of bulge and disk components of 135 <u>galaxies</u> from the CALIFA IFS survey.

According to Ph.D. student Iris Breda (IA & Science Faculty of the



University of Porto), what they found "implies that the formation timescale of bulges is inversely related to total galaxy mass: bulge formation in massive galaxies is completed within the first 4 billion years of cosmic evolution, whereas it is still ongoing at a low pace in less massive ones."



Image of the large spiral galaxy NGC 1232. The colours of the different regions



are well visible: the central areas contain older stars of reddish colour, while the spiral arms are populated by young, blue stars and many star-forming regions. Credit: ESO

This study, published in *Astronomy & Astrophysics*, reveals a coherent new scenario for the formation of galaxy bulges. Breda adds: "our study reveals a clear continuity in the properties of bulges, which strongly argues against the standard picture of two opposite bulge formation scenarios. Instead, bulge growth is driven by a superposition of quickearly with slow-secular processes, the relative importance of which is regulated by the mass and density of galaxies."

Another goal of this project has been to assess the role of Active Galactic Nuclei (AGN), powered by matter accretion onto <u>super-massive</u> <u>black holes</u>. They found that AGNs are the dominant source of gas ionization in massive bulges, whereas negligible in younger low-mass bulges. This could have far-reaching implications for our understanding of the co-evolution between bulges and super-massive black holes.

To FCT researcher Polychronis Papaderos (IA & University of Porto), "Our results are consistent with the notion both that the radiative efficiency of matter accretion onto Super-massive black holes scales with SMBH-mass or with a positive correlation between galaxy mass and super-massive black holes-to-bulge mass ratio. A further exploration of these hypotheses is of considerable interest."

This computationally demanding project went beyond previous ones, with regard to the amount of data analyzed, by the fact that it provided an accurate separation of bulge and disk, but also because, for the first time, a post-processing of the inferred star formation histories with RemoveYoung was made. This way it has been possible to explore how



bulge and disk were formed.

The study found that the luminosity contribution of stars younger than 9 billion years tightly correlates with stellar mass, stellar surface density, age and level of chemical enrichment of galaxy bulges. This quantity is therefore a powerful new diagnostic of the physical and evolutionary properties of galaxy bulges.

To IA coordinator José Afonso (IA & Science Faculty of the University of Lisbon): "This is a wonderful demonstration of IA's scientific and technical capability to understand one of the biggest mysteries in Astrophysics – how galaxies were formed throughout the entire history of the Universe. The use of the efficient computational tools developed by IA's researchers, together with observations from some of the most powerful telescopes and instruments available, is opening up a new view, and a new understanding, on the assembly history of galaxies, not only across space but also time."

More information: Iris Breda et al. The continuous rise of bulges out of galactic disks, *Astronomy & Astrophysics* (2017). DOI: 10.1051/0004-6361/201731705

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