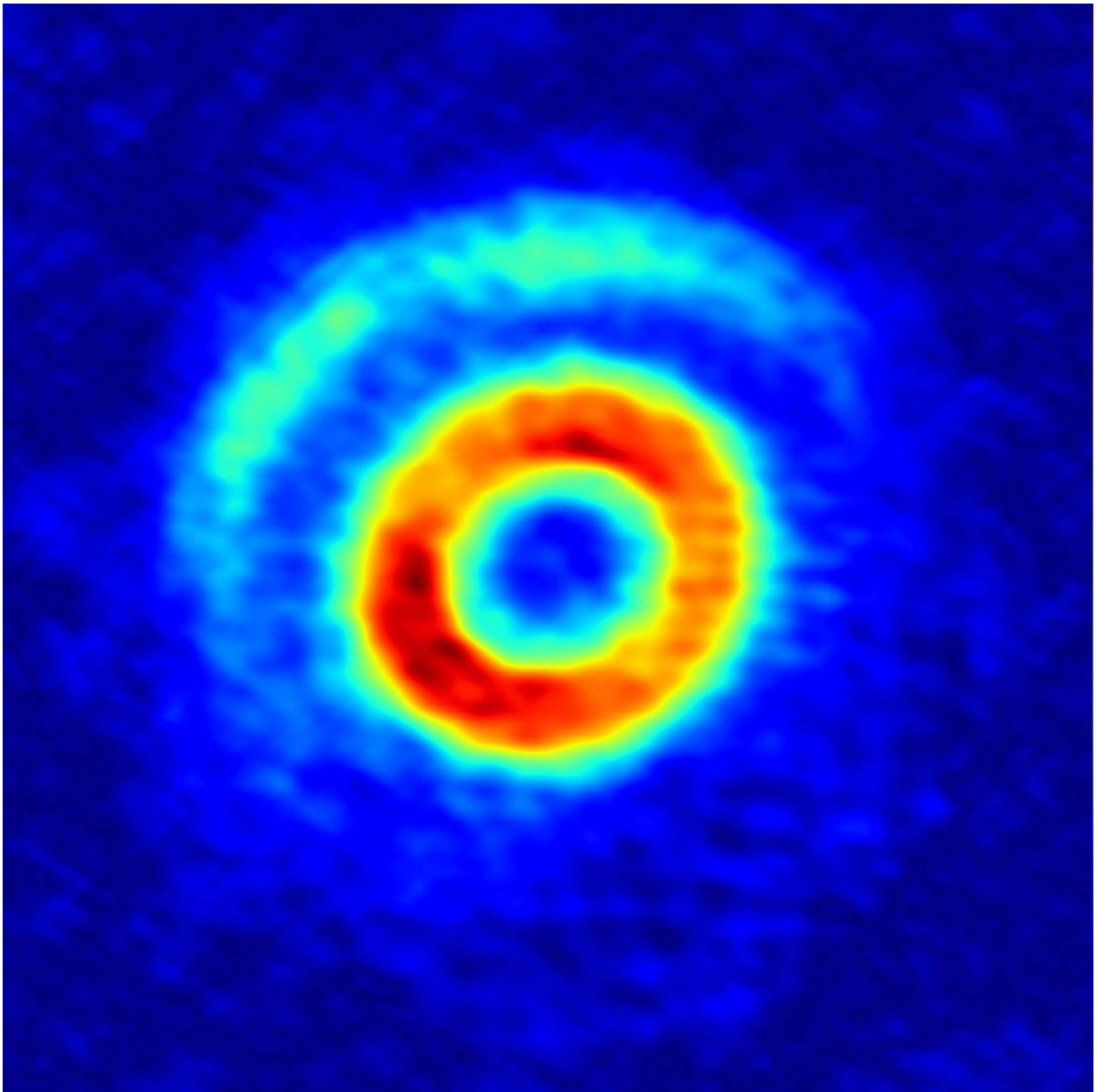


Astronomers find potential solution into how planets form

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V1247 Orionis, a young, hot star surrounded by a dynamic ring of gas and dust, known as a circumstellar disc. The disc can be seen in two parts: a clearly defined central ring of matter and a more delicate crescent structure located further out. Credit: Stefan Kraus

The quest to discover how planets found in the far reaches of the universe are born has taken a new, crucial twist.

A new study by an international team of scientists, led by Stefan Kraus from the University of Exeter, has given a fascinating new insight into one of the most respected theories of how [planets](#) are formed.

Young [stars](#) start out with a massive disk of gas and dust that over time, astronomers think, either diffuses away or coalesces into planets and asteroids.

However, scientists are still searching for a complete understanding of how these early formations come together to form asteroid-sized objects. One reason has been that drag in the disk produced by surrounding gas makes the grains move inward toward the star - which can in turn deplete the disk rapidly in a process known as "radial drift."

In the new research, the team use high powered telescopes to target the star V1247 Orionis -, a young, hot star surrounded by a dynamic ring of gas and dust.

The team produced a detailed image of the star and its surrounding [dust disc](#), shown in two parts: a clearly defined central ring of matter and a more delicate crescent structure located further out.

The region between the ring and crescent, visible as a dark strip, is

thought to be caused by a young planet carving its way through the disc. As the planet moves around in its orbit, its motion creates areas of high pressure on either side of its path, similar to how a ship creates bow waves as it cuts through water.

These areas of [high pressure](#) could become protective barriers around sites of planet formation; [dust particles](#) are trapped within them for millions of years, allowing them the time and space to clump together and grow.

Professor Kraus said: "The exquisite resolution of ALMA allowed us to study the intricate structure of such a dust-trapping vortex for the first time. The crescent in the image constitutes a dust trap that formed at the outer edge of the dark strip.

"It also reveals regions of excess dust within the ring, possibly indicating a second [dust trap](#) that formed inside of the putative planet's orbit. This confirms earlier computer simulations that predicted that [dust](#) traps should form both at the outer edge and inner edge of disc gaps.

"Dust trapping is one potential solution to a major stumbling block in our theories of how planets form, which predicts that particles should drift into the central star and be destroyed before they have time to grow to planetesimal sizes."

Dust-trapping vortices and a potentially planet-triggered spiral wake in the pre-transitional disk of V1247 Orionis is published in *Astrophysical Journal Letters*.

More information: Stefan Kraus et al. Dust-trapping Vortices and a Potentially Planet-triggered Spiral Wake in the Pre-transitional Disk of V1247 Orionis, *The Astrophysical Journal* (2017). DOI: 10.3847/2041-8213/aa8edc , [iopscience.iop.org/article/10. ...](https://iopscience.iop.org/article/10.3847/2041-8213/aa8edc)

[847/2041-8213/aa8edc](https://phys.org/news/2017-10-astronomers-potential-solution-planets.html)

Provided by University of Exeter

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