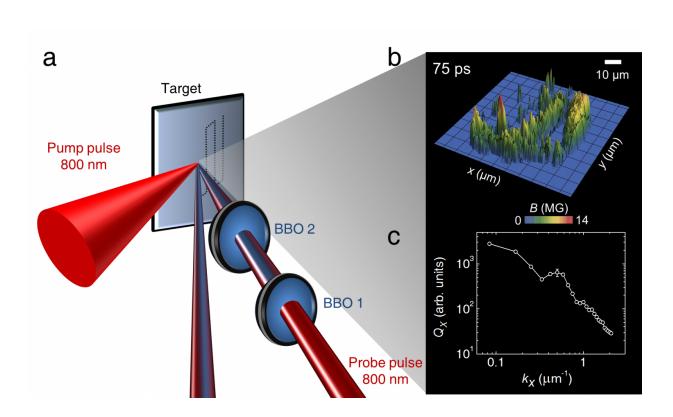


## Table top plasma gets wind of solar turbulence

June 30 2017



(a) A pump pulse creates the plasma on a solid while a probe pulse monitors the time evolution. The top panel in b shows the spatially randomized magnetic field in the plasma while the figure at the bottom shows the power spectrum of the magnetic field (magnetic energy density variation of with inverse of spatial length. The curve indicates magnetic turbulence in the plasma and mimics that from the solar wind. Credit: G. Chatterjee et al., *Nature Communications*, 10.1038/NCOMMS15970



Scientists from India and Portugal recreated solar turbulence on a tabletop using a high intensity ultrashort laser pulse to excite a hot, dense plasma and followed the evolution of the giant magnetic field generated by the plasma dynamics. This opens the possibility of studying astrophysical phenomena like the evolution of stars, in the lab.

Turbulent magnetic field dynamics that explain astrophysical phenomena like the evolution of stars have thus far been obtained only through observations via telescopes and satellites. Now, a team of scientists from India and Portugal have recreated such magnetic turbulence on a tabletop in the lab, using a <u>high intensity ultrashort laser pulse</u> to excite a hot, <u>dense plasma</u> on a solid surface and followed the extremely fast evolution of the giant magnetic field generated by the plasma dynamics. This groundbreaking study will be published in *Nature Communications* on 30 June.

Turbulence is everywhere—from tea cups to tokamaks and from water jets to weather systems, it is something we all see and experience. Yet, even after centuries of serious scientific study, fluid turbulence is still not properly understood. While it is difficult to define turbulence simply, it has many recognizable features, the most common being the fluctuations in parameters like velocity and pressure, indicating randomization of the flow.

Turbulence is not all bad and destructive, despite phenomena like <u>air</u> <u>turbulence</u> on a flight during bad weather. One good feature is that it enables much faster mixing than possible with only normal, slow diffusion. For instance, the sugar added to a cup of tea would take hours or days to disperse undisturbed, but stirring makes the tea turbulent, resulting in fast mixing at the molecular level. Turbulence also helps in mixing fuel and oxygen for efficient combustion in engines.

Much of our universe consists of highly ionized gas known as plasma,



which can often be extremely hot and swirling at unimaginable speeds. Turbulence in a plasma is much more complex than that in neutral hydrodynamic fluids. In a charged plasma environment, the negatively charged, light electrons and positive heavy ions respond at vastly different length and time scales. The motion of these charged species is governed by electromagnetic forces and the current flow through the charge particle dynamics leads to magnetic field generation. Therefore, the randomness of magnetic fields often mimics the fluid turbulence in plasmas.

The team of scientists leading this new study find that the turbulence in the magnetic <u>field</u> is initially driven by the electrons (at a trillionth of a second) and the ions step in and take over at longer times. This is the first time such a 'relay race' involving two different species has been glimpsed. Further, these lab observations have an uncanny resemblance to the satellite data on the <u>magnetic field</u> spectra measured for turbulent astrophysical plasmas in the solar wind, solar photosphere and Earth's magnetosheath. Although in the laser experiment the electrons in the plasma are initially energised, the ion dominant response that kicks in at later times shows spectral features similar to those in the astro systems. These experiments thus establish clear connections between the two scenarios, even though the driver of <u>turbulence</u> in the lab <u>plasma</u> is very different from that in the astrophysical system.

More information: *Nature Communications* (2017). DOI: <u>10.1038/NCOMMS15970</u>

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