

Scientists observe first Sun-like magnetic cycle on another star

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An international team of scientists led by the University of Göttingen has observed a Sun-like magnetic cycle on another star for the first time. The Sun's magnetic field drives the Sun's spots and flares and fuels the Solar wind – a torrent of material that streams off our star into space. The discovery is important not only for stellar physics, but also to understand and predict how the Sun affects the Earth and our technological society through its magnetic activity. The results were published in *Astronomy & Astrophysics*.

With the advent of dedicated instruments known as stellar spectropolarimeters roughly ten years ago, it became possible to map the magnetic fields of nearby Sun-like <u>stars</u>. Using this new technology at the Bernhard Lyot Telescope in the French Pyrenees, the scientists observed the star 61 Cyg A over a period of nine years. Lying in the northern constellation of Cygnus, 61 Cyg A is somewhat smaller and less massive than the Sun, and at a distance of just over eleven light years it is one of the Sun's nearest neighbours.

The Sun's activity varies over the course of a 22-year long magnetic cycle, with the polarity of its magnetic field flipping every eleven years. The frequency and strength of these activities wax and wane over the course of a cycle, with two active periods interspersed with more quiet ones. All in all, the variations are relatively small and slow – a stark contrast to the great bulk of known magnetically active stars that vary dramatically in brightness, release enormous <u>flares</u> and display much more complex long-term variability.



Although 61 Cyg A is a little dimmer and cooler than the Sun, the <u>scientists</u> were able to detect changes in its activity coinciding with polarity flips over a seven-year activity cycle, for a magnetic cycle of 14 years. They observed polarity changes every seven years and an increased complexity in its <u>magnetic field</u> when these flips were approached.

"Our findings could contribute greatly towards creating models of how the Sun and other stars generate magnetic fields. This will enable us to gain an understanding of this important process, which is thought to be operating inside all Sun-like stars, and to help us to further understand our own Sun," explains Sudeshna Boro Saikia, Ph.D. student at Göttingen University and lead author of the study. A better understanding of this process and of our Sun in general will increase our ability to predict the impact of the Sun's activities on our technology on Earth and on orbiting satellites.

The Solar wind and coronal mass ejections can indeed have a huge impact on Earth. When these flows of plasma reach Earth, they not only produce the northern and southern lights, but they can also disturb radio communication and power grids at ground level, as well as damage satellites and even threaten astronauts in Earth orbit.

More information: S. Boro Saikia et al. A solar-like magnetic cycle on the mature K-dwarf 61 Cygni A (HD 201091), *Astronomy & Astrophysics* (2016). DOI: 10.1051/0004-6361/201628262

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