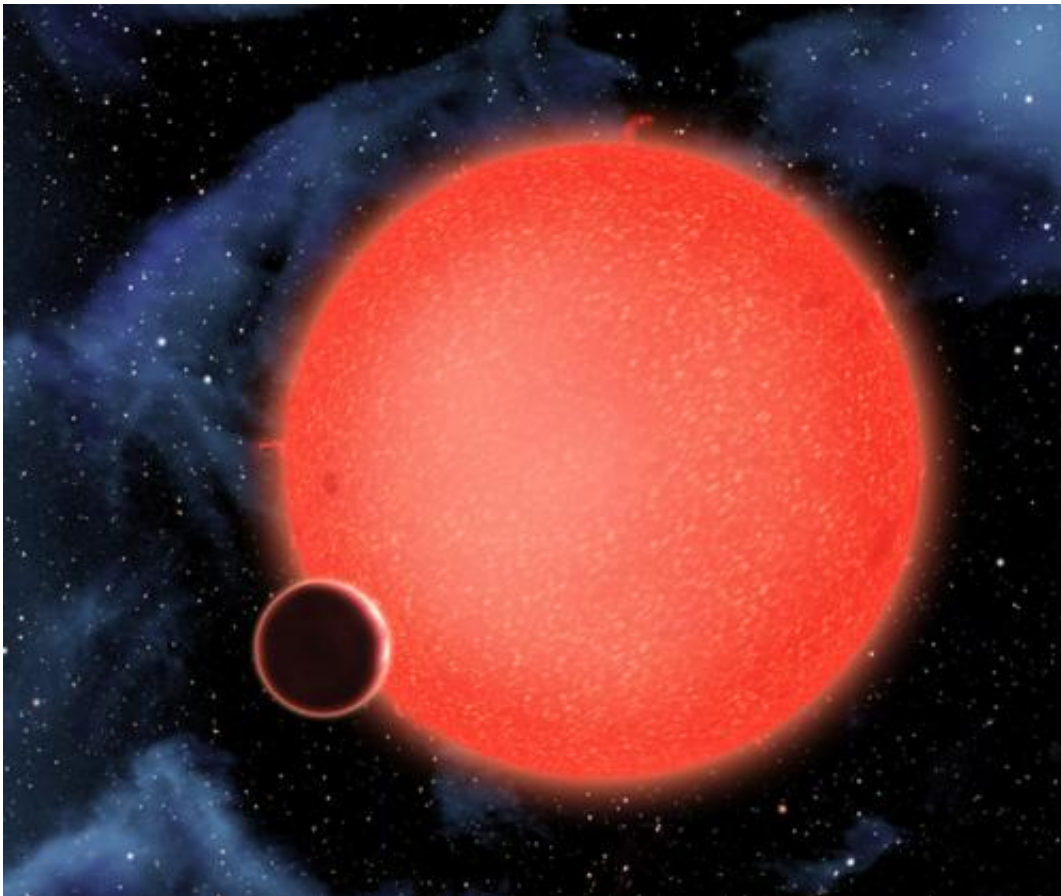


# Water-world Earths could host life, even if they're askew

February 26 2015, by Elizabeth Howell, Astrobiology Magazine

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Artist's conception of GJ1214b, a super-Earth that could have a surface dominated by ocean, orbiting its red dwarf star. Observations with the Hubble Space Telescope revealed a thick atmosphere. Credit: NASA, ESA, and D. Aguilar (Harvard-Smithsonian Center for Astrophysics)

Life could be habitable on an Earth-sized waterworld tilted on its side if the oceans aren't too shallow, a new study reveals.

As long as the entire world was covered in oceans of at least 50 meters (165 feet) deep, temperatures would be moderate enough at the poles to support life. Even at the equator, which would be the chilliest part of that world since it only would receive a bit of sunlight in spring and fall, life could still exist.

But if you were to shrink the ocean's depth to something like 20 meters (66 feet), then the risk of a runaway cold effect becomes much greater. Should a thin veneer of ice develop in the ocean, it's possible the [climate](#) system would collapse into an ice block in just a few hundred years. That short timeline would be tough for life to develop a foothold.

"That's a bad outcome for life," said lead author David Ferreira, who was with the Massachusetts Institute of Technology when the study was conducted. "With deeper oceans, a collapse into a global snowball is possible, but a bit harder. It feeds into the idea that if you have an extensive, big, deep ocean, your chances to find life or a climate that is habitable are higher."

Ferreira's paper, titled "Climate at high-obliquity," was published in the journal *Icarus* in November. It forms part of his greater research interest in the role of oceans on climate. That research interest led Ferreira to the United Kingdom's University of Reading, where he is a lecturer (professor) in the department of meteorology.

## Hot poles, cold equator

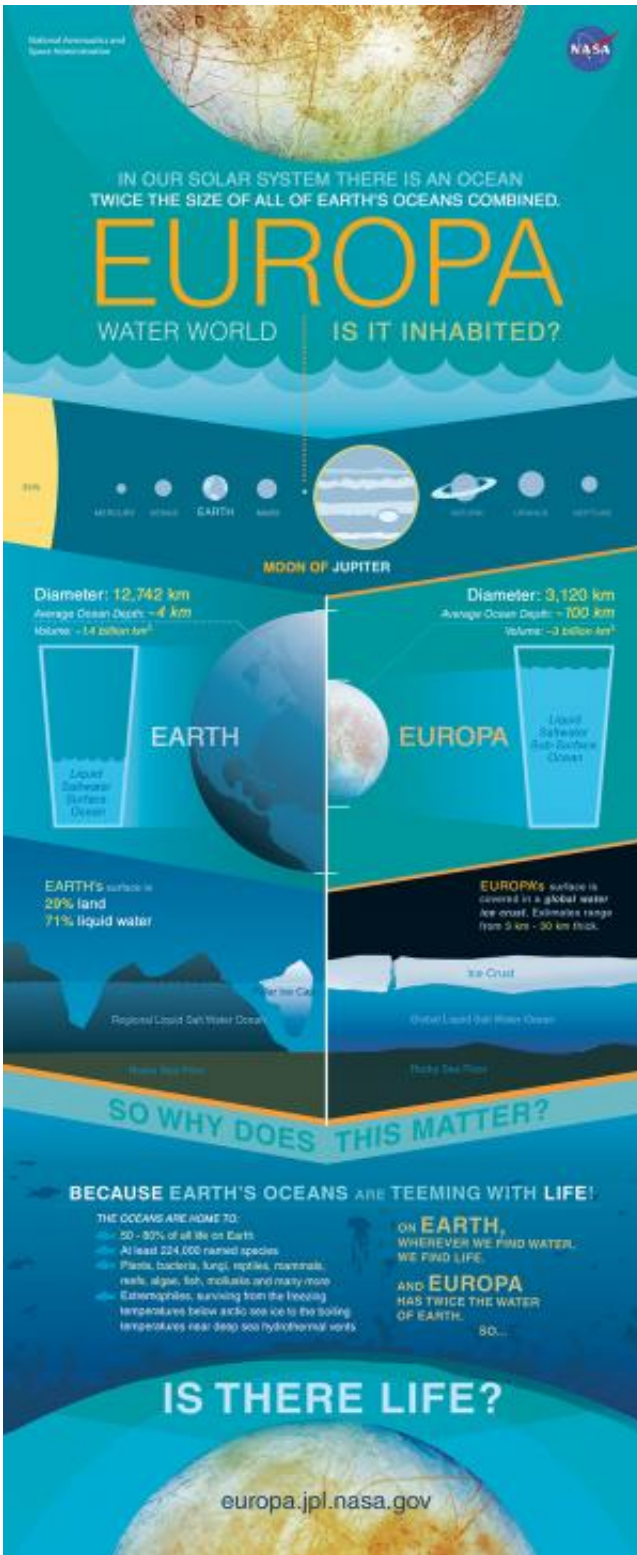
The traditional view of "habitable" [planets](#) came from looking at those that are in the "Goldilocks zone" of their parent stars. This is the point where water can exist above the freezing point, but it's not so hot that the

water begins to boil away.

The conditions for life, however, are more complicated than that. For example, if the planet is too large, the pressure of the gas will likely make it too tough for life to survive. If the planet is too small, its gravity could be too low to hold on to an atmosphere. Therefore, many researchers say [habitable planets](#) in the Goldilocks zone must be close to Earth's size.

Other factors can also come into play, such as the presence or absence of an ocean. As those who live in coastal California or the south of Italy know, the proximity of water can make temperatures over nearby land much more steady and mild. On a planet-size scale, a [global ocean](#) would also do this trick as long as it is deep enough, the research reveals.

For simplicity's sake, the simulation assumed an Earth-sized planet orbiting a sun-like star at the same distance our planet does (93 million miles, or 150 million kilometers). The researchers, however, changed two major parameters.



An infographic of Europa, an icy “waterworld” moon of Jupiter that could also be host to life. Credit: NASA Jet Propulsion Laboratory

The first was the planet's tilt. Earth's axis is tilted at 23.5 degrees, which produces enough of a difference across the planet to produce the seasons. The simulation instead made the tilt 90 degrees so that the planet was spinning on its side.

The second variable was the presence of oceans. While the Earth is covered in oceans by about 70 percent, the simulation assumed 100 percent cover with different depths, ranging from 10 meters (33 feet) to about 3,000 meters (657 feet). It was the threshold of 50 meters that interested researchers the most, as this was considered a minimum depth to have a stable climate suitable for life.

The poles would seem to be the toughest place to live on this theoretical world. During the summer, they would face the sun directly, while in the winter they would face away. Even in the coldest part of the year, the surface temperature in those zones would be no less than 10 to 15 degrees Celsius (50 to 59 degrees Fahrenheit.)

"It's a bit like the Earth's Arctic in the summer," Ferreira said.

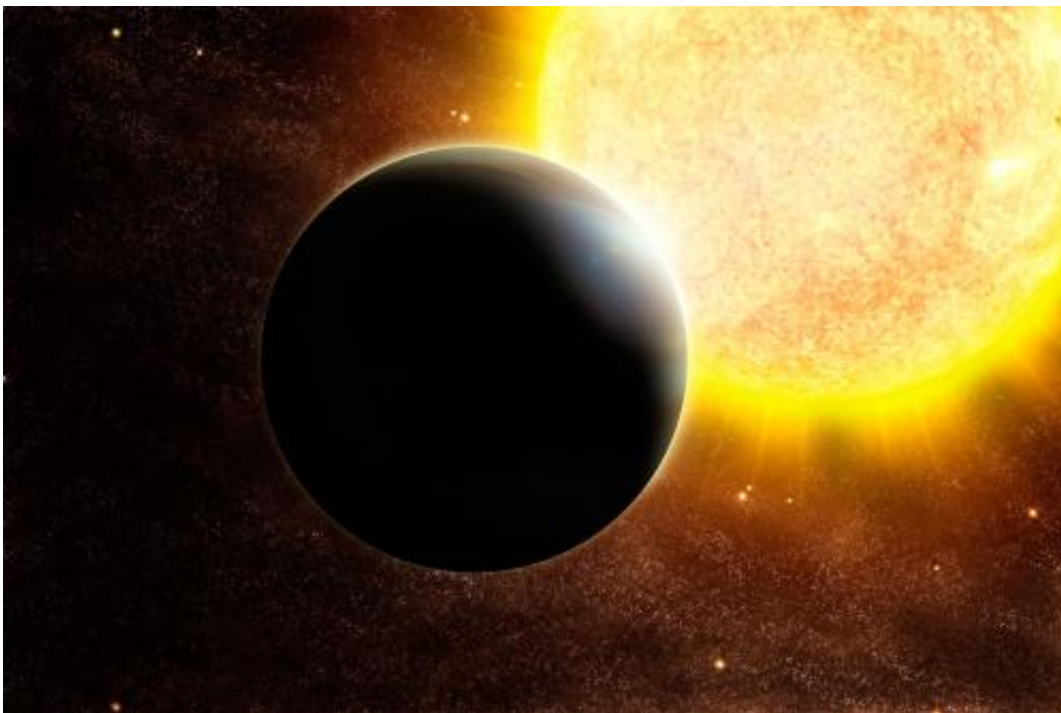
The summer, by contrast, would see temperatures soar to 35 to 40° Celsius (95 to 104° Celsius). That's hot, but by no means hot enough to discourage life from surviving. Meanwhile, the equators would be the coldest parts of the planet, but would remain above freezing, at 2 to 4° Celsius (36 to 39° Fahrenheit.)

"Even there, those are not harsh conditions. Liquid water would survive there," Ferreira pointed out.

While waves were not simulated on this water world—they're too small for the scale of the simulation—what was examined was the role of thermal currents. The researchers found similar current systems to Earth's, which are driven by temperature differences in the ocean and

atmospheric winds. There is, for example, a well-known circulation pattern on Earth that brings water from the Southern Hemisphere to the North Atlantic.

"It's typical of what people would do with climate simulations for future global warming. It's on this level of complexity," Ferreira said.



Planets close in to their stars, such as this Jupiter-sized one in an artist's illustration, are more likely to be detected. Credit: ESO

## Mapping for future planet-hunters

There are other kinds of worlds where habitability could be possible, in the case of a global [ocean](#). Other systems ripe for consideration include "super-Earths"—those planets that are slightly larger than our own—and "mini-Neptunes," or planets that are a bit smaller than the gas-swaddled

Outer Solar System planet.

What the researchers are considering next, however, is a "tidally locked" planet. This is a planet that perpetually has one side facing its star, and another facing away. This kind of configuration is common in our own solar system. Earth's moon is tidally locked to our planet. Jupiter and Saturn also have small moons (relative to the gas giants' size) that keep one side facing the planet.

It's too early to make predictions as to how habitable those worlds could be, but Ferreira said if habitability is possible, this increases researchers' chances of finding life beyond the Solar System. Tidally locked worlds are actually among the easiest kinds of exoplanets for researchers to find. This is because of the methods astronomers use to seek out new worlds. One of them relies on measuring the gravitational "wobble" a planet produces on its parent star. If the planet is closer to its star, it will have a stronger pull, which makes it easier to be detected.

Another method looks for the disc of a planet passing across the disc of its star. Planets with close-in orbits would make those crossings more frequently than planets that don't, which again increases the odds of them being detected with current technology.

Earth-sized worlds, however, are hard to find due to their tiny size. That said, NASA's Kepler space telescope has detected at least two in the habitable regions of their parent stars. Future telescopes could make the search easier, since they could be more sensitive to smaller planets.

Upcoming planet-hunters include NASA's James Webb Space Telescope (slated for launch in 2018), and the candidate European mission, PLATO (PLANetary Transits and Oscillations of stars), which would launch in 2024.

Ferreira's research, however, will continue in the direction of oceans on

newfound worlds.

"Oceans on the Earth are the big regulator of the climate system," he said. "Naturally, the question is how you would apply that knowledge to the planets that are in a different astronomical state than Earth. One would expect oceans in such planets would be a strong regulator on the climate as well, and a factor in habitability."

**More information:** David Ferreira, John Marshall, Paul A. O'Gorman, Sara Seager, "Climate at high-obliquity," *Icarus*, Volume 243, 15 November 2014, Pages 236-248, ISSN 0019-1035, [DOI: 10.1016/j.icarus.2014.09.015](https://doi.org/10.1016/j.icarus.2014.09.015).

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