

## Winter in the southern uplands of Mars

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Perspective view of Hooke crater in Argyre basin. Credit: ESA/DLR/FU Berlin

Over billions of years, the southern uplands of Mars have been pockmarked by numerous impact features, which are often so closely packed that they overlap. One such feature is Hooke crater, shown in this frost-tinged scene, imaged by ESA's Mars Express during winter in the southern hemisphere.

Hooke <u>crater</u> is located near the northern edge of the 1800 km-wide Argyre basin – one of the most impressive impact structures on Mars, excavated in a giant collision about 4 billion years ago.



Sitting in a flat part of the basin known as Argyre Planitia, Hooke crater has a diameter of 138 km and a maximum depth of about 2.4 km. It is named after the English physicist and astronomer Robert Hooke (1635–1703).

Hooke crater comprises two different impact structures, with a smaller impactor blasting a depression off-centre in the floor of a larger, preexisting crater. The newer crater in the centre is filled with a large mound topped by a dark dune field. The mound appears to be composed of layered material, possibly alternating sheets of sand and frost.

Dark dunes also spread southwards (to the left in the topographic, main colour and 3D images) from the smaller crater, partially covering the floor of the main crater. The local topography modifies the airflow, serving as a sand trap for the wind-blown sediments.

In these images, much of the low-lying region to the south, as well as the central mound inside Hooke crater, is covered with a thin, white coating of carbon dioxide frost. At higher elevations and on north-facing crater walls, the frost is largely absent and appears only in areas shaded from direct sunlight by the walls of smaller craters.





Hooke crater is located at about 46°S / 316°E, near the northern edge of the 1800 km-wide Argyre basin, one of the most impressive impact structures on Mars. The crater, with diameter 138 km and a maximum depth of about 2.4 km, is situated in a flat region within the basin known as Argyre Planitia. Credit: NASA MGS MOLA Science Team

Outside the crater, there are a number of linear features visible on the floor of Argyre Planitia, on the south (left) side of the topographic, main colour and 3D images. These are examples of 'yardangs', rocky ridges that have been shaped by prolonged wind erosion. Most of the yardangs are oriented towards Hooke crater, indicating the prevailing wind direction.



Also visible on the floor of the Argyre basin are small areas of chaotic terrain, which resemble depressions containing flat-topped mesas, buttes and hills. In the topographic, main colour and 3D images, one of these regions can be seen at the top edge, about a third of the way from the left, and another in the lower middle part, down from the left-most edge of the crater.

Chaotic terrains like these are thought to have been created when largescale melting of ground ice caused the ground to collapse. Where the terrain has not collapsed completely, the larger mesas may still contain substantial water ice.



Colour-coded topography map of a region of the Argyre basin, featuring Hooke crater and part of the floor of the basin known as Argyre Planitia. White and red show the highest terrains, while blue and purple show the deepest. The image is based on a digital terrain model of the region, from which the topography of the landscape can be derived. The region clearly slopes to the south (left). Credit: ESA/DLR/FU Berlin





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Data from the nadir channel and one stereo channel of the High Resolution Stereo Camera on Mars Express have been combined to produce this anaglyph 3D image, which can be viewed using stereoscopic glasses with red–green or red–blue filters. Credit: ESA/DLR/FU Berlin

## Provided by European Space Agency

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