

## Europe's billion-star surveyor set for launch

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Gaia is an ambitious mission to chart a three-dimensional map of our Galaxy, the Milky Way, in the process revealing the composition, formation and evolution of the Galaxy. Gaia will provide unprecedented positional and radial velocity measurements with the accuracies needed to produce a stereoscopic and kinematic census of about one billion stars in our Galaxy and throughout the Local Group. This amounts to about 1 per cent of the Galactic stellar population. Combined with astrophysical information for each star, provided by on-board multi-colour photometry, these data will have the precision necessary to quantify the early formation, and subsequent dynamical, chemical and star formation evolution of the Milky Way Galaxy. Additional scientific products include



detection and orbital classification of tens of thousands of extra-solar planetary systems, a comprehensive survey of objects ranging from huge numbers of minor bodies in our Solar System, through galaxies in the nearby Universe, to some 500 000 distant quasars. It will also provide a number of stringent new tests of general relativity and cosmology. Gaia will be launched from Europe's Spaceport in French Guiana at the end of 2013. Credit: ESA–D. Ducros, 2013

By repeatedly observing a billion stars, with its billion-pixel video camera, the Gaia mission will allow astronomers to determine the origin and evolution of our galaxy whilst also testing gravity, mapping our inner solar system, and uncovering tens of thousands of previously unseen objects, including asteroids in our solar system, planets around nearby stars, and supernovae in other galaxies.

Professor Gerry Gilmore, from the University of Cambridge and UK Principal Investigator for Gaia, said, "Gaia will be a revolution in our knowledge of the local Universe. For the first time we will have a fair sample of what is out there, where it is, how it is moving, how unseen (dark) matter is distributed, where and when stars formed and where and when the chemical elements of which we are made were created. Gaia will make a huge step towards understanding how the Milky Way came to be formed, and evolved into what we see today. For the first time, we will be able to see the Milky Way in 3-D. In fact in 6-D – where stars are, and how they are moving."

UK participation in the mission is funded by the UK Space Agency and scientists and engineers from around the UK have played key roles in the design and build of Gaia. The UK Science and Technology Facilities Council (STFC) funded the early development of the project, including the set-up of the data applications centre. STFC's current support involves the UK exploitation of the scientific data to be yielded from the



mission.

The Cambridge Gaia Data Processing Centre will be the front line in processing Gaia's images, which will also be key to the discovery of many thousands of transient stars and supernovae: these will be made immediately available to schools and the public for their participation in the research.

Dr Chris Castelli, Acting Director of Technology, Science and Exploration at the UK Space Agency, said, "Gaia is an important space mission for the UK; we've won around €80 million of contracts from the European Space Agency to build the spacecraft and are providing a state of the art data centre that will turn the mission's raw data into the largest stellar catalogue ever made."

Andy Stroomer, Astrium UK's Director of Earth Observation, Navigation, and Science, added, "We are extremely proud of our contribution to the unique Gaia mission - the latest example in a long and successful heritage of supporting ESA science. Stevenage engineers have provided core systems for the Gaia satellite including video processing unit, satellite electrical platform, and mechanical subsystems,".

Gaia will map the stars from an orbit around the Sun, near a location some 1.5 million km beyond Earth's orbit known as the L2 Lagrangian point. The spacecraft will spin slowly, sweeping its two telescopes across the entire sky and focusing their light simultaneously onto a single digital camera, the largest ever flown in space.

Once Gaia starts routine operations, around Easter 2014, astronomers will have the challenge of dealing with a flood of data. Even after being compressed by software, the data produced by the five-year mission will fill over 30 000 CD ROMs. This data will be transmitted 'raw' and will need processing on Earth to turn it into a calibrated set of measurements



that can be freely used by the astronomical community. The cutting edge computer technology developed at the Cambridge Data Processing Centre will be key to this process.

The first Gaia science will be discoveries of new sources – supernovae, extreme variable stars and blazars - which will be discovered at the Cambridge processing centre, and immediately made available for study by both professionals and the interested public. Gaia will discover many new sources which are bright enough for amateurs, and schools with access to public robotic telescopes, to become the first to confirm and obtain more information. Schools and amateurs will be able to load their data onto the Gaia web site (gaia.ac.uk" target="\_blank">gaia.ac.uk) where it will be used in scientific analyses, and fully credited.

"A school class can be the first to `adopt a supernova', observe it with robotic telescopes, such as the Faulkes, and provide critically important science information which we need to understand the new sources. That class will be doing original science, and will be credited for their research contributions," says Dr Heather Campbell, a scientist at Cambridge who is part of the Gaia science alerts analysis team.

"By participating in the Gaia Alerts programme, and remotely controlling the Faulkes Telescopes to observe exciting targets discovered just hours beforehand, UK schools will be making real contributions to the research side of this project." added Professor Paul Roche, Director of the FT project at the University of South Wales.

## The UK and Gaia

The UK has two major roles in the Gaia mission: building the spacecraft, and delivering the science.

UK industry and science institutes won some €80 million of industrial



contracts to build Gaia, with leadership roles in building the heart of Gaia, the array of 106 CCDs, the control avionics and the critical micropropulsion system, as well as playing a critical role in the development of the Gaia spectrometer.

Astrium at Stevenage was responsible for the spacecraft's super precision guidance and control system as well as the powerful on-board computers needed to process the torrent of data it will produce.

The 'eye' of Gaia's camera has the most sensitive set of light detectors ever assembled for a space mission. It is also the largest focal plane array ever to be flown in space and contains a mosaic of 106 large area, high performance Charged Coupled Device (CCD) CCD91-72 image sensors, which are custom designed, manufactured and tested by UK company e2v. Without these image sensors the Gaia mission would not be possible. These detectors were calibrated with the Gaia electronics at Mullard Space Science Laboratory.

SciSys UK Ltd is responsible for the spacecraft's operational simulator.

Gaia data will be processed and analysed ready for release to the scientific community and public at six data centres, including one in the UK, operating software developed and tested by a consortium of 400 people across Europe, including some 50 people at 6 Institutes in the UK (Cambridge, UCL-MSSL, Leicester, Edinburgh, The Open University, STFC RAL Space and Bristol).

University College London's Mullard Space Science Laboratory (MSSL) has a major role in spectroscopic science, and in conjunction with The Open University, is involved in software development including architecture, integration and validation, preprocessing development, spectra extraction and calibration.



The UK hosts the Cambridge data processing centre, where the Gaia imaging data are processed. These data provide brightness, colour, and position information, which enable the core Gaia science, determining distances motions and the intrinsic properties of each star. Gaia also has a spectrograph, to measure the radial velocities of 300 million stars, delivering their complete 3-D positions and 3-D velocities. The spectra will be processed at the French data centre, using software in part developed in the UK.

## Details

Shortly after its launch from Europe's space port in French Guiana, Gaia will deploy its sunshield, forming a 10.5 m-wide 'skirt' around the spacecraft's base. The shield has two purposes: to shade Gaia's sensitive telescopes and cameras from sunlight, allowing them to cool to their operating temperature of  $-110^{\circ}$ C, and to provide power to operate the spacecraft. The Sun-facing side of the shield is partially covered with solar panels to generate electricity.

Gaia will be placed in an orbit around the Sun, at the second Lagrange point L2, which is named after its discoverer, Joseph Louis Lagrange (1736-1813). For the Sun-Earth system, the L2 point lies at a distance of 1.5 million kilometres from the Earth in the anti-Sun direction and corotates with the Earth in its 1-year orbit around the Sun. An operational lifetime of 5 years is planned.

The Gaia spacecraft will be controlled from the European Space Operations Centre (ESOC, Darmstadt, Germany) using the two ground stations Cebreros (Spain) and Perth (Australia).

Science operations will be conducted from the European Space Astronomy Centre (ESAC, Villafranca, Spain).



Mission Operations Commissioning will take place during the transfer phase of the mission, when the spacecraft is flying towards its destination: the second Lagrange point of the Sun-Earth system (L2). This early commissioning phase deals with the Service Module and with all Payload Module functions that can be checked-out during the transfer phase. For this purpose, during the transfer phase the nominal attitude with respect to the Sun is the same as the operational one. The final commissioning of the Payload Module occurs after orbit insertion around L2, consisting of final calibration of the instrument and demonstration of instrument performance.

Provided by European Space Agency

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