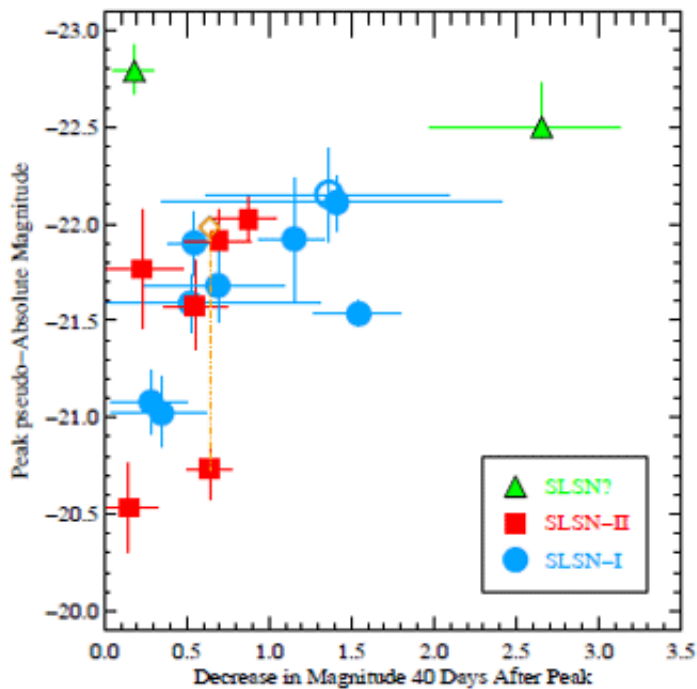


How often do stars explode as exceptionally bright supernovae?

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It is not yet clear what gives rise to supernovae with a peak brightness many times the average, so-called superluminous supernovae (SLSNe), but since they are observable from further away than normal supernovae, a better understanding of these events might also make them a valuable additional 'standard candle' for distance measurements in the Universe. Using survey data from the ROTSE-IIIb telescope at the McDonald

Observatory (Texas, US) that covered 500 square degrees of sky, a small team of international researchers – including CAASTRO member Dr Fang Yuan (ANU) – has now calculated the volumetric rate of SLSNe.

Having identified five suitable events, the first step in their calculations was to produce light curve templates and estimate pseudo-absolute magnitude distributions for both hydrogen-poor SLSN Type I and hydrogen-rich SLSN Type II. Monte Carlo simulations were then used to determine the efficiency of different surveys on the telescope in 'shortlisting' candidates in a given volume of sky. For the pooled SLSN-like data and at a [redshift](#) of 0.2 (local volume), the team calculated a volumetric rate of 199 events $\text{Gpc}^{-3} \text{yr}^{-1} h^3$ which, due to the small sample size, is subject to large statistical (199 +137 / -86) and systematic (199 +65 / -41) errors.

Their results approximately match the local rate of sub-energetic, long-duration gamma-ray bursts but are exceeded by the estimated rate of core collapse supernovae by a factor of 400 to 1300. These new calculations now offer an opportunity to determine the origin of SLSNe by comparing them with the formation rate of stars in the [critical mass](#) range at similar redshifts. And since peak magnitudes of SLSN Type I were found to be tightly clustered ($M = -21.7 \pm 0.4$), these events might be a promising 'standard candle' once sufficient sample sizes have been reached.

More information: Quimby, R. et al. Wheeler in MNRAS 431 "Rates of Superluminous Supernovae at $z \sim 0.2$ " arxiv.org/abs/1302.0911

Provided by CAASTRO

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