

Hot cores in dark clouds

December 6 2011



A false-color, infrared image of a dark cloud as seen against the bright background of Milky Way emission. Within the snake-like cloud structure seen here, dust and gas are congealing into new stars that are detected as embedded, hot cores. New observations with the Submillimeter Array find that these hot cores are often rich in organic chemicals. Credit: NASA/JPL-Caltech

(PhysOrg.com) -- The earliest stages in the life of a star are among the most mysterious. This is primarily because stars form inside dark clouds of material that block optical light, and because they form relatively quickly, in only hundred of thousands of years, whereas once a star starts burning its hydrogen fuel it can last for billions of years.

The details of high [mass star](#) birth are the most difficult of all to unravel because massive stars evolve the fastest, and can reach maturity even before their natal birth cloud dissipates. As a result, astronomers infer their presence from hot cores buried deep inside dark clouds, perhaps

accompanied by other signposts like outflowing gas or strong [ultraviolet radiation](#).

Infrared dark clouds -- dark patches in the sky seen against the bright infrared galactic background emission -- are rich in molecules and dense gas. They are natural sites for [star birth](#), and the hot cores inside them are thought to be among the earliest stellar embryos. CfA astronomers Steve Longmore and Qizhou Zhang, together with four colleagues, used the Smithsonian's [Submillimeter Array](#) (SMA) to examine two hot cores that were previously discovered in infrared surveys and then confirmed as cores at millimeter wavelengths.

The SMA observations revealed the presence of over a dozen complex molecules including complex nitrogen-, oxygen-, and sulfur-rich species and isotopes of these molecules. The main molecules are CH₃CN, CH₃OH, CH₃OCH₃, CH₃OCHO, SO, SO₂, NH₂CHO, C₂H₅OH, C₂H₅CN, CN, NH₂CHO, H₂CS, OCS, and HNCO, although the rich, densely packed spectrum makes some of the identifications uncertain. The SMA provides maps of the hot cores in each of these molecular species. The scientists model the data to conclude that the masses of the two cores are about eight and twenty-six solar-masses, that the temperatures in the material near these young stars are a few hundred kelvin, and that there is evidence for rotating structures (possibly preplanetary disks) around the young stars. The result helps to confirm that these hot cores are massive stars, and the rich chemistry helps trace the details of the stellar birth process in its very early phases.

Provided by Harvard-Smithsonian Center for Astrophysics

Citation: Hot cores in dark clouds (2011, December 6) retrieved 20 April 2024 from <https://phys.org/news/2011-12-hot-cores-dark-clouds.html>

This document is subject to copyright. Apart from any fair dealing for the purpose of private study or research, no part may be reproduced without the written permission. The content is provided for information purposes only.