

# Young clays on Mars could have been habitable regions for life

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(PhysOrg.com) -- Two small depressions on Mars found to be rich in minerals that formed by water could have been places for life relatively recently in the planet's history, according to a new paper in the journal *Geology*.

“We discovered locations at Noctis Labyrinthus that show many kinds of minerals that formed by [water](#) activity,” said Catherine Weitz, lead author and senior scientist at the Planetary Science Institute. “The clays we found, called iron/magnesium (Fe/Mg)-smectites, are much younger at Noctis Labyrinthus relative to those found in the ancient rocks on [Mars](#), which indicates a different water environment in these depressions relative to what was happening elsewhere on Mars.”

Smectites are a specific type of clay mineral that readily expands and contracts with adsorbed water. They contain Silica, plus Aluminum, Iron or Magnesium in their structures. They form by the alteration of other silicate minerals in the presence of non-acidic water.

Weitz and her co-authors studied approximately 300 meters of vertically exposed layered rocks within two 30 to 40 kilometer depressions, called troughs, near the western end of the Valles Marineris canyon system. Using high-resolution images from the High Resolution Imaging Science Experiment (HiRISE) camera and hyperspectral data from the Compact Reconnaissance Imaging Spectrometer for Mars (CRISM) on the Mars Reconnaissance Orbiter (MRO) spacecraft, combined with Digital Terrain Models (DTMs) to determine elevations and view geometric relationships between units, the team was able to map hydrated minerals and understand how the water chemistry varied with time within each trough, said Weitz, a HiRISE team member.

Each trough probably experienced multiple episodes where water partially filled in low-lying regions and deposited minerals. As each trough continued to enlarge and experience collapse over time, older minerals became buried and separated, followed by deposition of younger minerals, then finally erosion to re-expose buried units. Volcanism from the Tharsis volcanoes to the west may have created subsurface water that was subsequently transported through the ground and into the troughs. Localized volcanism that produced ash and gases, hydrothermal activity, and melting snow/ice within the troughs could have also produced some of the minerals. The observed minerals indicate water varied in pH levels over time, in one trough from acidic to neutral, and in the other trough from neutral to acidic and back to neutral.

Other occurrences of Fe/Mg-smectites have been found on Mars but almost exclusively in association with older, Noachian-age (more than

3.6 billion years ago) rocks, or produced by younger impact events. Following the deposition of Fe/Mg-smectites in the Noachian period, the climate on Mars is believed to have changed during the Hesperian time to favor formation of minerals under more acidic conditions, such as salts rich in sulfur (sulfates).

Weitz and her co-authors identified the same sulfates and Fe/Mg-smectites in the Noctis Labyrinthus troughs found elsewhere on Mars, but the progression of minerals over time, from sulfates to Fe/Mg-smectites, indicates a reverse order relative to what happened globally across Mars.

“These clays formed from persistent water in neutral to basic conditions around 2 to 3 billion years ago, indicating these two troughs are unique and could have been a more habitable region on Mars at a time when drier conditions dominated the surface,” said co-author and CRISM team member Janice Bishop from the SETI Institute and NASA AMES Research Center.

“These troughs would be fantastic places to send a rover, but unfortunately the rugged terrain makes it unsafe both for landing and for driving,” Weitz said.

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Provided by Planetary Science Institute

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