

Dusty experiments are solving interstellar water mystery

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Molecular cloud. Credit: NASA/JPL-Caltech/L. Allen (Harvard-Smithsonian CfA)

(PhysOrg.com) -- Dust may be a nuisance around the house but it plays a vital role in the formation of the key ingredient for life on Earth - water - according to researchers at Heriot-Watt University. The results from pioneering experiments to solve one of the mysteries of the interstellar space, where did all the water come from, will be presented by Victoria Frankland at the RAS National Astronomy Meeting in Glasgow.

“We think that the Earth’s water was delivered by comets during the early stages of Earth’s history and that comets were formed from interstellar material left over after the birth of the Sun, but the next step back has been unclear,” said Ms Frankland.

Water is relatively abundant in the interstellar medium and hydrogen atoms are extremely common, but there is a problem with the other vital ingredient for H₂O. Gas phase reactions that can take place in the interstellar medium are limited by the low temperatures and pressures. Experiments show that it is possible for [hydrogen atoms](#) to combine with molecules of oxygen (O₂) or ozone (O₃) under the conditions of the interstellar medium. However, observations by recent satellite missions have detected very little gaseous molecular oxygen (O₂) and [ozone](#) (O₃) has never been detected at all in these regions of space. On the other hand, atomic oxygen (O) is quite plentiful, but gas phase reactions between hydrogen and atomic oxygen can't account for the amount of water observed. Even the observed quantities of atomic oxygen suggest that some is 'missing' in star-forming regions compared to the rest of interstellar space.

Ms Frankland and her colleagues at Heriot-Watt believe the [dust grains](#), which make up about 1% of the [interstellar medium](#), hold the key by providing a surface that helps reactions take place. In addition, some molecules remain stuck to the surface, building up an icy coating over time. This coating, which is mainly [water ice](#), can then play a role in reactions.

The team at Heriot-Watt has pioneered surface science techniques to evaluate experimentally exactly how such reactions might occur. However, the temperatures in [interstellar space](#) can reach just a few degrees above absolute zero, so recreating the conditions in the laboratory has been a challenge.

“Our experiments rely on being able to reproduce in the laboratory the very low pressures and low temperatures of these star-forming regions. We set up our experiments in a vacuum chamber and cool it down to -268 degrees Celsius, then use surface sensitive techniques to explore the physical and chemical behaviour of oxygen atoms and molecules on the

surfaces of dust and ice grains.” said Ms Frankland.

Initially, the experiments have been looking at how the surfaces of dust particles affect the reactions of oxygen in its various forms in order to eliminate other water formation reactions. However, the ultimate aim of this research will be to combine atomic beams of oxygen and hydrogen study in situ water formation on a grain surface.

“These initial experiments are having some interesting results in that they are allowing us to look at how the ice coating develops on the dust particles. It appears that oxygen atoms may become trapped inside the icy mantles. We need to do more work, but it may be that our experiments might help solve the mystery of the missing [atomic oxygen](#) as well as where the water has come from,” said Ms Frankland.

Provided by Royal Astronomical Society

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